



Physics of Sawtooth and its control

Valentin Igochine

Max-Planck Institut für Plasmaphysik
EURATOM-Association
D-85748 Garching bei München
Germany

Introduction

- Motivation

- Stability of (1,1) mode in tokamak

Sawtooth period and amplitude (slow time scales)

- Sawtooth model for crash prediction

- Influence of the fast particles

- Control of fast particle stabilized sawteeth

- New control approaches

Crash event (fast time scales)

- Experimental observations

- Possible models

- Stochastic model

- Computation results

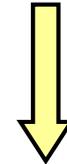
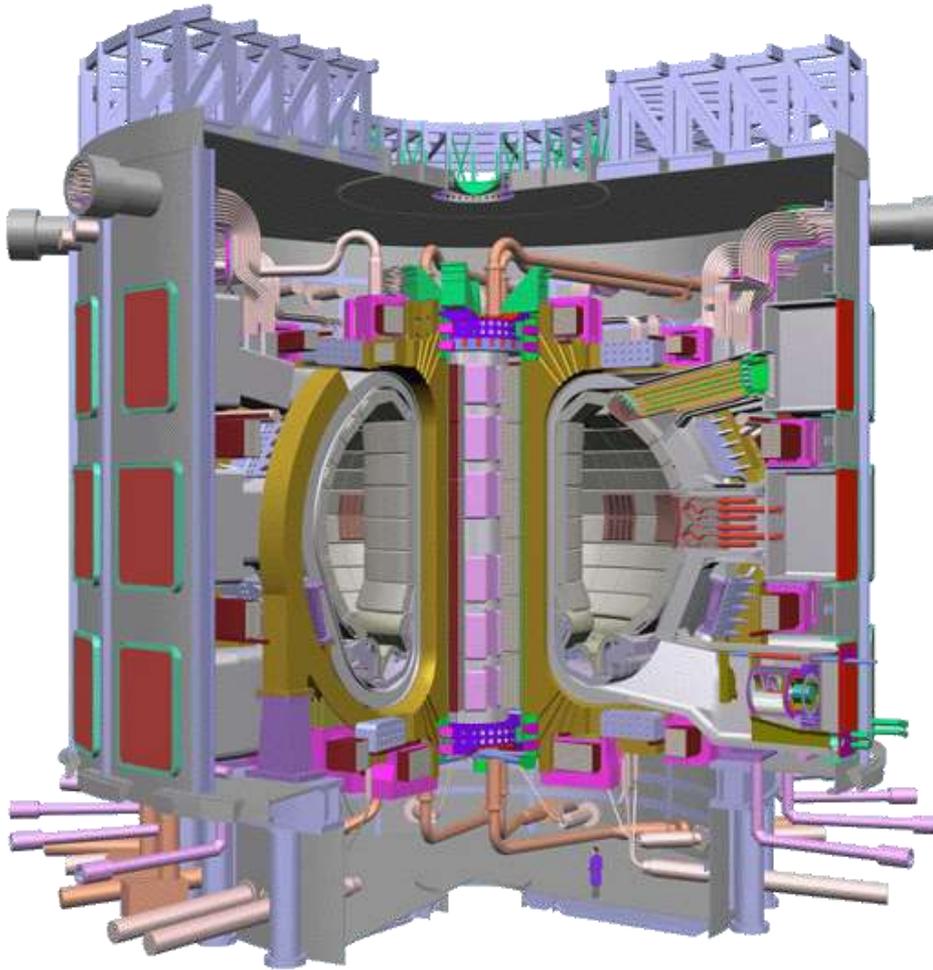
Conclusion

Why do we need to control sawteeth?

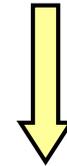


- **Problem with long Sawteeth?**

ITER will have large fusion-born alpha particle population in the core



Long sawtooth periods



Trigger NTMs and cause confinement degradation (reduced fusion power... etc), or worse disruption



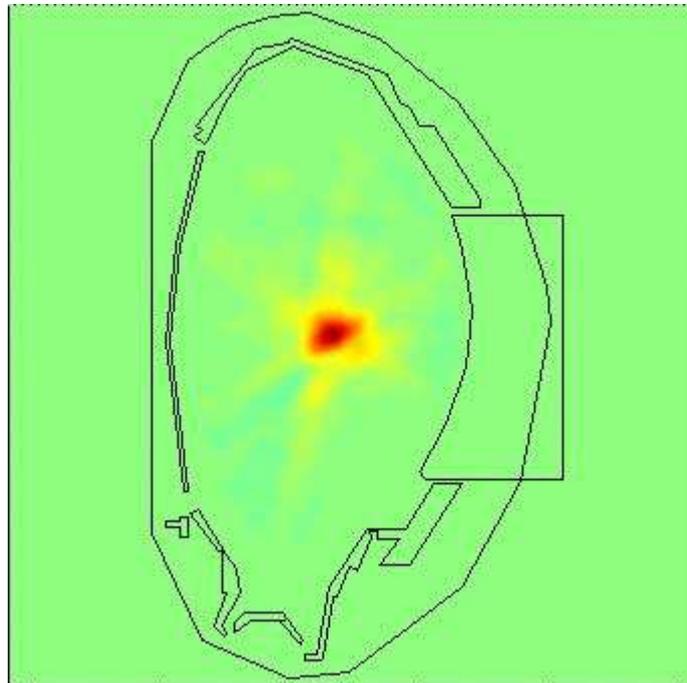
Since NTMs can be at large amplitude very rapidly, a prevention approach recommended: keep sawteeth small and frequent

SXR tomography in ASDEX Upgrade

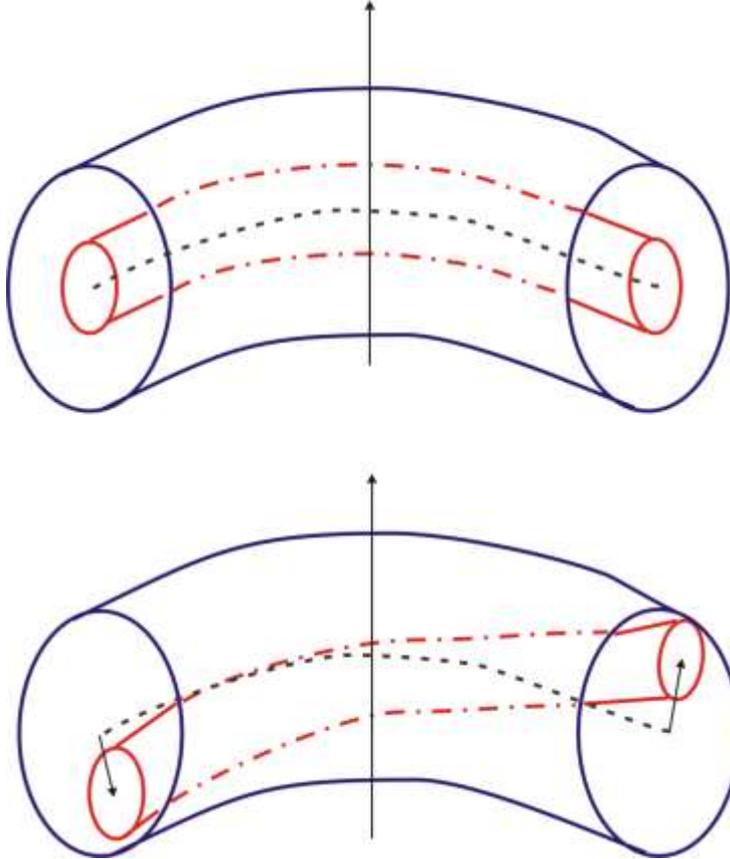


Topographic reconstruction of sawtooth crash in ASDEX Upgrade from soft X-ray cameras signals (maximal entropy method).

Discharge: 23074 $t=3.07s$ (color scale is different for each time frame)

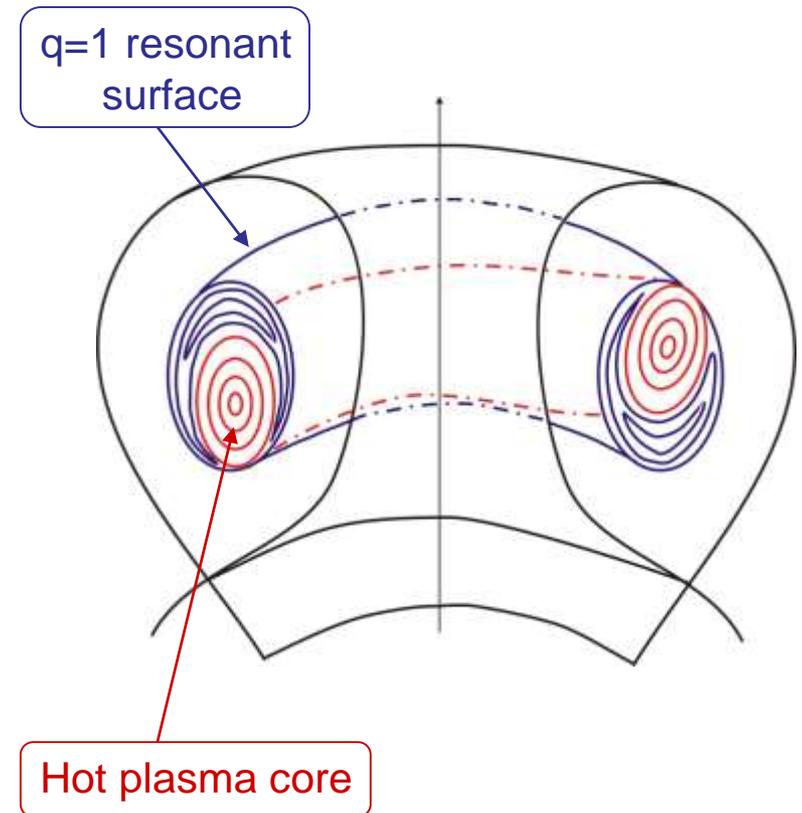


Internal Kink instability in a Tokamak.



A Tilt and Shift of the Core Plasma.

Sawteeth: internal (1,1) kink mode.



Energy principle: $\delta W(\zeta, \xi) = -\frac{1}{2} \int \zeta \cdot \mathbf{F}(\xi) d^3x$

Low β approximation, cylinder (good for current driven modes):

$$\mathbf{j}_\perp = \frac{\mathbf{B} \times \nabla p}{B^2} = \mathcal{O}(\varepsilon^2), \quad \varepsilon = \frac{a}{R_0} \ll 1$$

$$\mathbf{B} = B \begin{pmatrix} 0 \\ \frac{r}{qR_0} \\ 1 \end{pmatrix} + \mathcal{O}(\varepsilon^2), \quad \mathbf{j} = \frac{B}{R} \begin{pmatrix} 0 \\ 0 \\ \frac{1}{r} \left(\frac{r^2}{q}\right)' \end{pmatrix} + \mathcal{O}(\varepsilon^2)$$

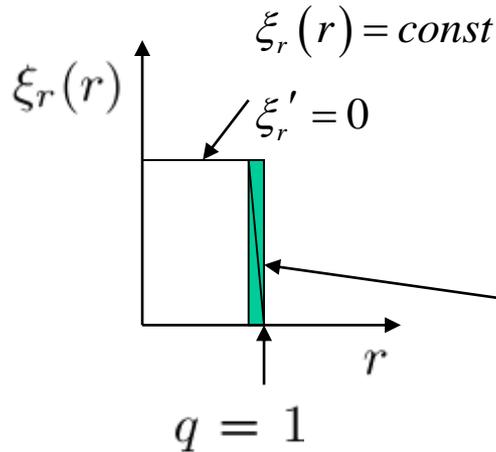
$$\xi = \xi(r) e^{i(m\theta - n\phi - \omega t)} \quad \xi = \xi_r \hat{\mathbf{r}} + \xi_\theta \hat{\boldsymbol{\theta}} + \xi_\parallel \mathbf{b}$$

$$\delta W_{\text{cyl}} = \pi^2 \frac{B_0^2}{R_0} \int_0^a (|r\xi_r'|^2 + (m^2 - 1)|\xi_r|^2) \left(\frac{n}{m} - \frac{1}{q}\right)^2 r dr + \mathcal{O}(\varepsilon^4)$$

(1,1) internal kink mode is a special case

$$\delta W_{\text{cyl}} = \pi^2 \frac{B_0^2}{R_0} \int_0^a (|r\xi_r'|^2 + (m^2 - 1)|\xi_r|^2) \left(\frac{n}{m} - \frac{1}{q}\right)^2 r dr + \mathcal{O}(\varepsilon^4)$$

for $m=1$
$$\delta W_{m=1} = \pi^2 \frac{B_0^2}{R_0} \int_0^a r^3 |\xi_r'|^2 \left(1 - \frac{1}{q}\right)^2 dr$$



This function minimize the functional

δ width of the layer where $\xi_r(r)$ changes

$$|\xi_r'|^2 = \mathcal{O}(1/\delta^2) \quad \left(1 - 1/q\right)^2 = \mathcal{O}(\delta^2) \quad \text{at } q \approx 1$$

Small and large terms cancel in the integrand. It follows that the $\delta W_{m=1}$ vanishes because the integration interval itself has width δ

(1,1) internal kink mode is pressure driven



To understand stability of the (1,1) mode we have to compute next terms in the energy functional expansion $\delta W = \mathcal{O}(\varepsilon^4)$

If one takes into account now new terms (pressure gradient, perpendicular current, toroidal curvature, compressibility)

$$\delta W \approx 6\pi^2 \frac{B_0^2 r_1^4}{R_0^3} |\xi_r(0)|^2 [1 - q(0)] [\beta_{\text{crit}}^2 - \beta_p^2(r_1)]$$

$$\beta_{\text{crit}}^2 = \frac{13}{144}$$

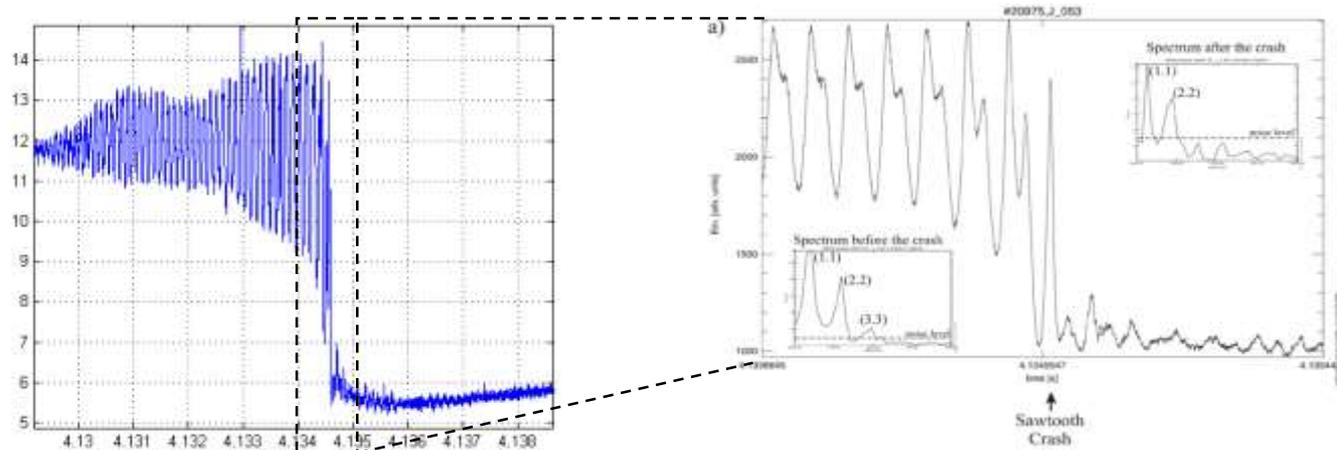
$$\beta_p(r_1) \equiv -2 \frac{R_0^2 q^2}{B_0^2 r_1^4} \int_0^{r_1} p' r^2 dr$$

Unstable only if $\beta_p > \beta_{\text{crit}}$!!
(1,1) internal kink mode is pressure driven!

This is different to other internal kinks which are current driven in the simplest approximation

M.N. Bussac, R. Pellat, D. Edery, and J.L. Soule, "Internal kink modes in toroidal plasma with circular cross-section," *Phys. Rev. Lett.* **35**, 1638 (1975).

Rem.: Energetical particles can compete with small MHD contribution and lead to strong interaction with particles (discussed later)



Understanding of the sawteeth

Slow time scale

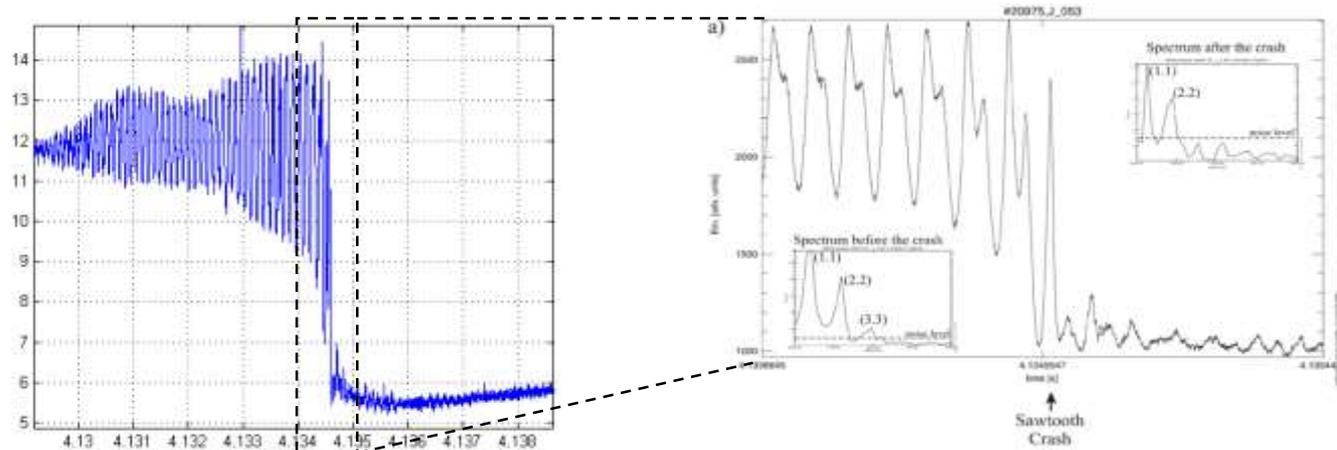
- sawteeth period
- mixing radius
- amplitude

Especially important for ITER

Fast time scale

- mechanism of the crash

Understanding of the basic physics



Understanding of the sawteeth

Slow time scale

- sawtooth period
- mixing radius
- amplitude

Especially important for ITER

Fast time scale

- mechanism of the crash

Understanding of the basic physics

Porcelli model for sawteeth (slow time scale)



Porcelli-Boucher-Rosenbluth model for sawtooth period and amplitude

[Porcelli et al, PPCF, 38, 2163 (1996)]

Linear stability of the (1,1) mode is considered.

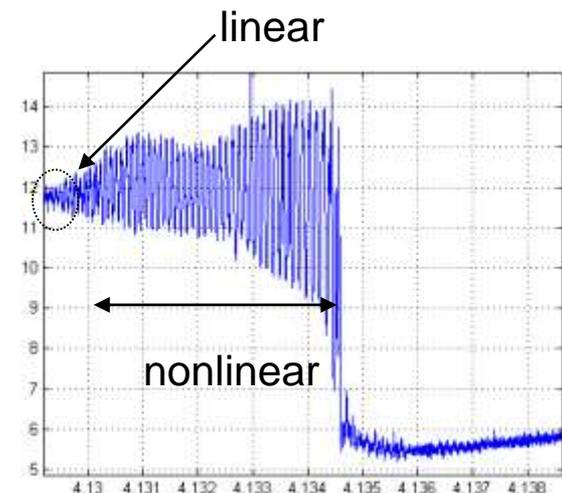
Sawtooth crashes are triggered by internal (1,1) kink modes if:

- $q(0) < 1$
- resistive (1,1) becomes unstable or ideal (1,1) becomes unstable

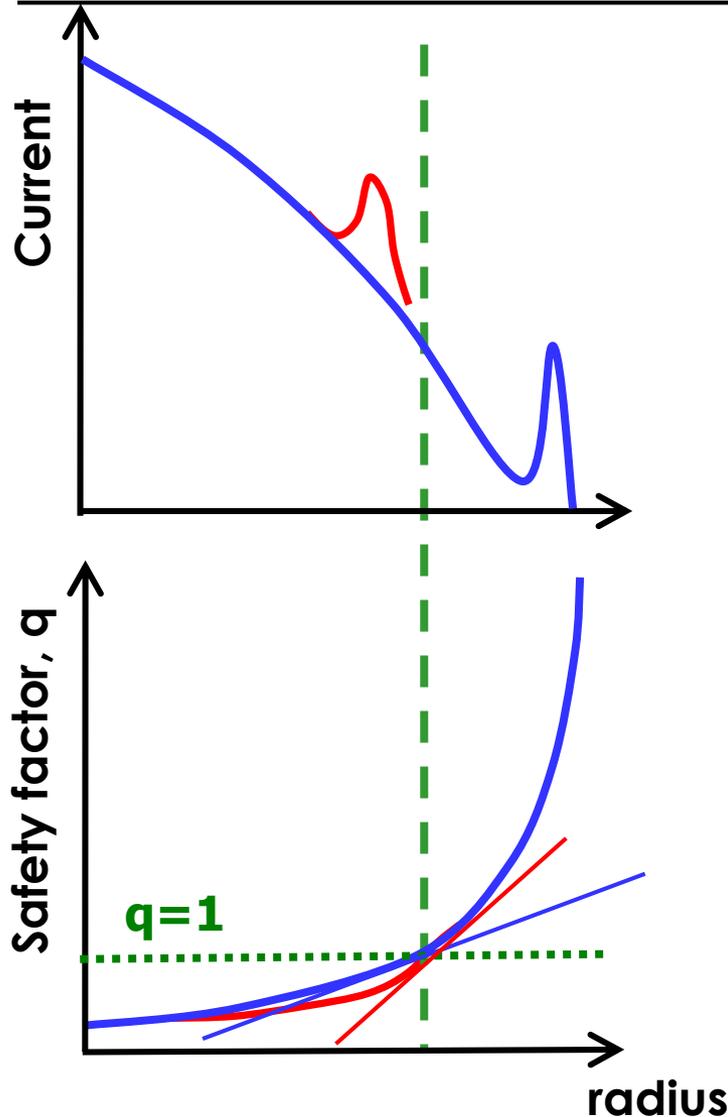
$$\frac{\delta \hat{W}}{s_1} < const$$

Model works in spite of the fact that (1,1) mode is strongly nonlinear !

(This is just a luck!)

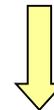


Driving current changes the sawtooth stability

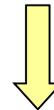


Trigger condition: $\pi \frac{\delta \hat{W}}{s_1} < c_\rho \frac{\rho}{r_1}$

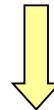
A perturbation is applied to the initial current profile (inside $q=1$)



More current means more poloidal field, so q drops



$q=1$ is moved outwards



Magnetic shear at $q=1$, s_1 (~gradient of q) increases

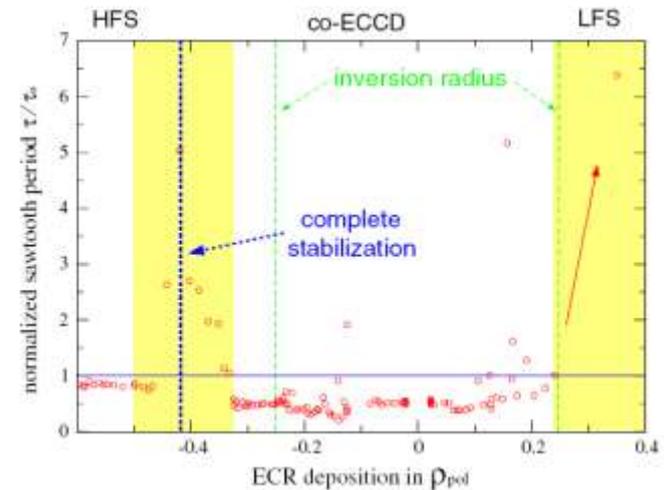
What else influence the sawtooth stability ?

$$\frac{\delta \hat{W}}{s_1} < const$$

Local current drive

Influence on the shear at $q=1$

Example: Change of shear at $q=1$ with co-ECCD and counter-ECCD in ASDEX Upgrade [A.Mueck, PPCF, 2005]



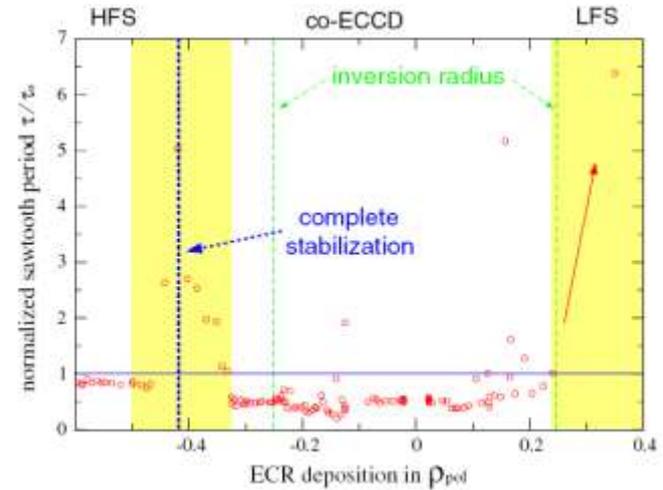
What else influence the sawteeth stability ?

$$\frac{\delta \hat{W}}{s_1} < const$$

Local current drive

Influence on the shear at $q=1$

Example: Change of shear at $q=1$ with co-ECCD and counter-ECCD in ASDEX Upgrade [A.Mueck, PPCF, 2005]



$$\frac{\delta \hat{W}}{s_1} < const$$

Influence on the mode stability δW

Toroidal rotation

Relatively good understood

Plasma shape

Heating deposition

Profiles

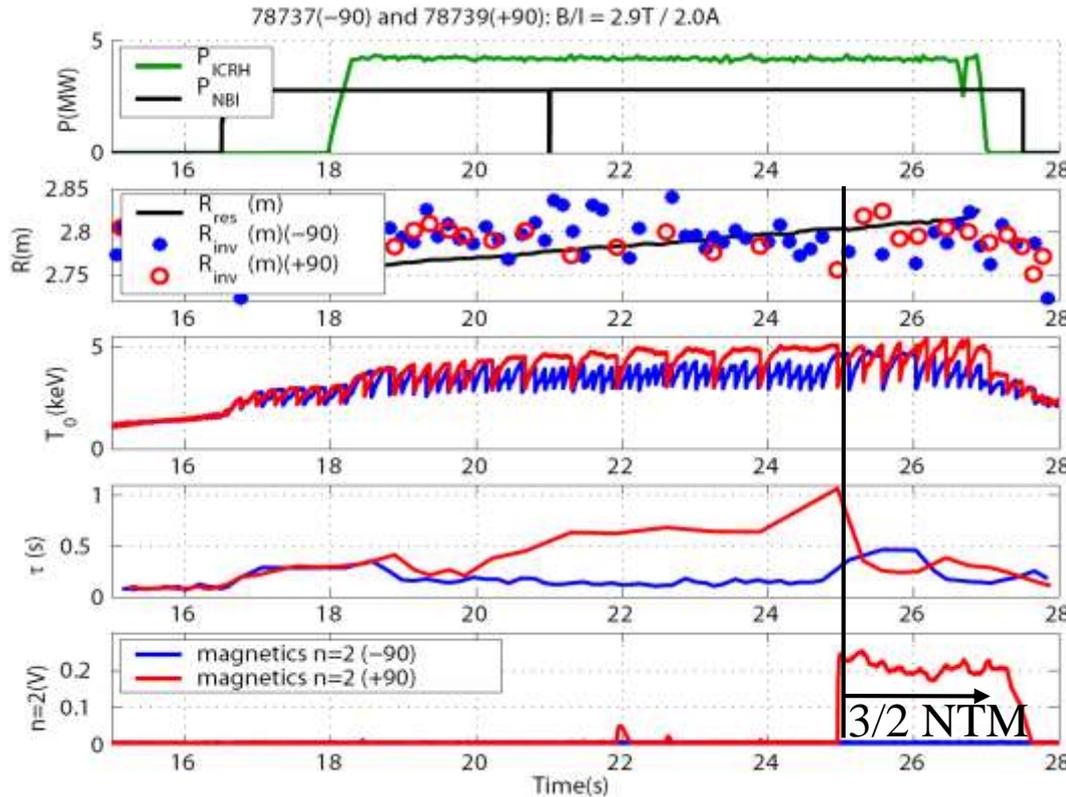
Fast particles

Fast particles are expected to be strongly stabilizing in ITER

Different ICRH heating of the ^3He minorities in JET



2.9T < B < 2.96T (co-, cuntr- propagating)



Current drive is negligible because we are influence on minority (no influence on s_1)

Main effect has kinetic nature (change $\delta\hat{W}$)

$$\frac{\delta\hat{W}}{s_1} < const$$

[Graves et al, NF 2010]

Different ICRH heating of the ^3He minorities in JET

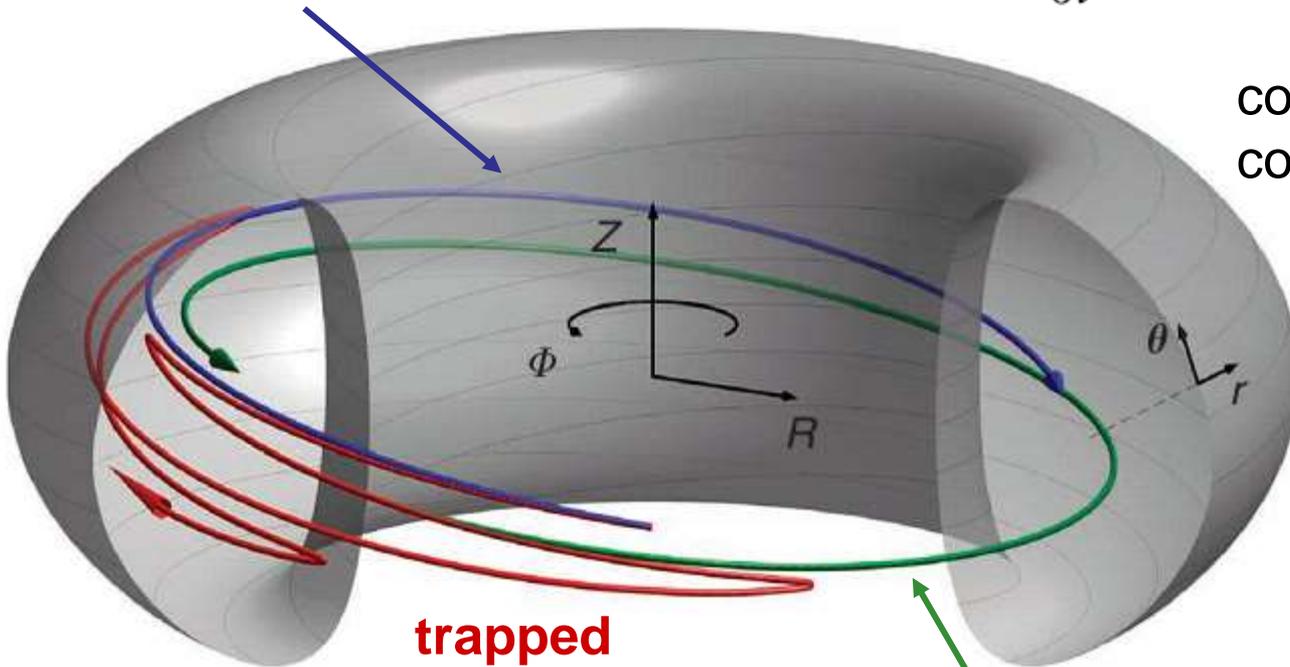


$$\rho \frac{\partial^2 \xi}{\partial t^2} = \delta \mathbf{J} \times \mathbf{B} + \mathbf{J} \times \delta \mathbf{B} - \nabla \delta P_c + \mathcal{F}_h(\delta P_h)$$

core
collisional

hot
collisionless

co-passing

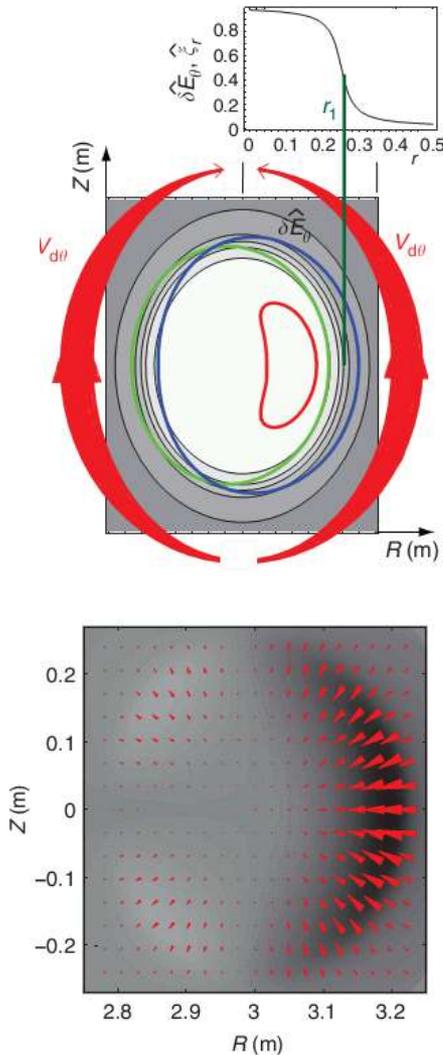


trapped

counter-passing

J.P. Graves, Nature, 2012

Different ICRH heating of the ^3He minorities in JET



$$\rho \frac{\partial^2 \xi}{\partial t^2} = \delta \mathbf{J} \times \mathbf{B} + \mathbf{J} \times \delta \mathbf{B} - \nabla \delta P_c + \mathcal{F}_h(\delta P_h)$$

core
(collisional)
force

hot ions
(collisionless)
force

Radial component of the normalized force points inward for trapped ions if displacement is outward. (Stabilizing effect from $\hat{\mathcal{F}}_h$!)

Physically this means conservation of third adiabatic invariant. (If low frequency perturbation is trying to change adiabatically the flux through these orbits, the orbits will tilt or shift in space in order to preserve this flux)

J.P. Graves, Nature, 2012

Different ICRH heating of the ^3He minorities in JET



Sawtooth Period (78737, 78739, 78740): $B / I = 2.9\text{T} / 2.0\text{A}$

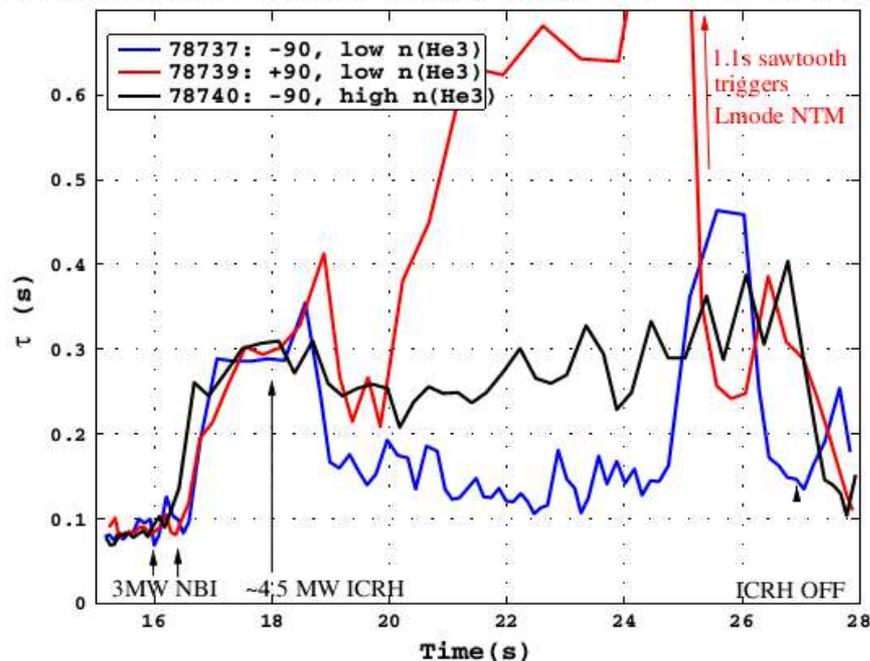


Figure 5. The sawtooth period for 78737 (-90° phasing, low concentration ^3He), 78740 (-90° phasing, high concentration ^3He) and 78739 ($+90^\circ$ phasing, shown also in figure 1).

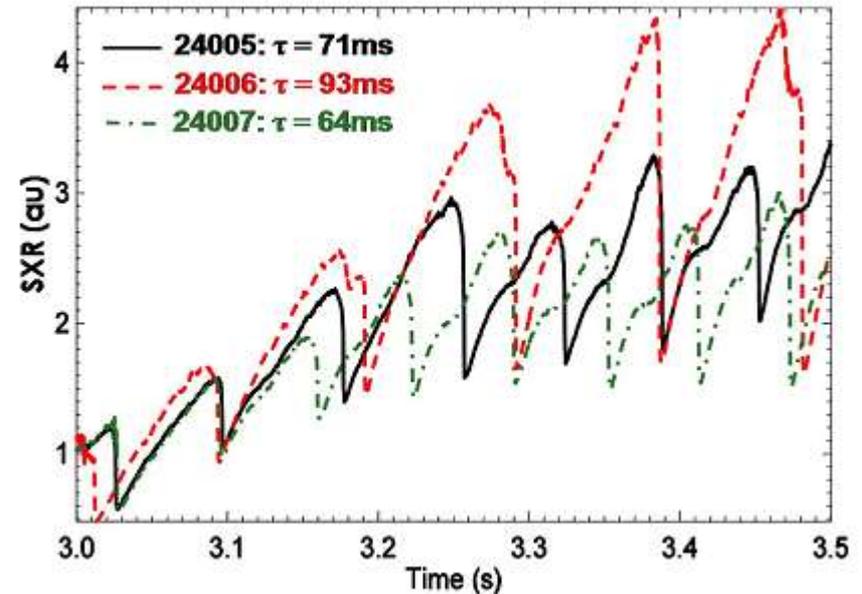
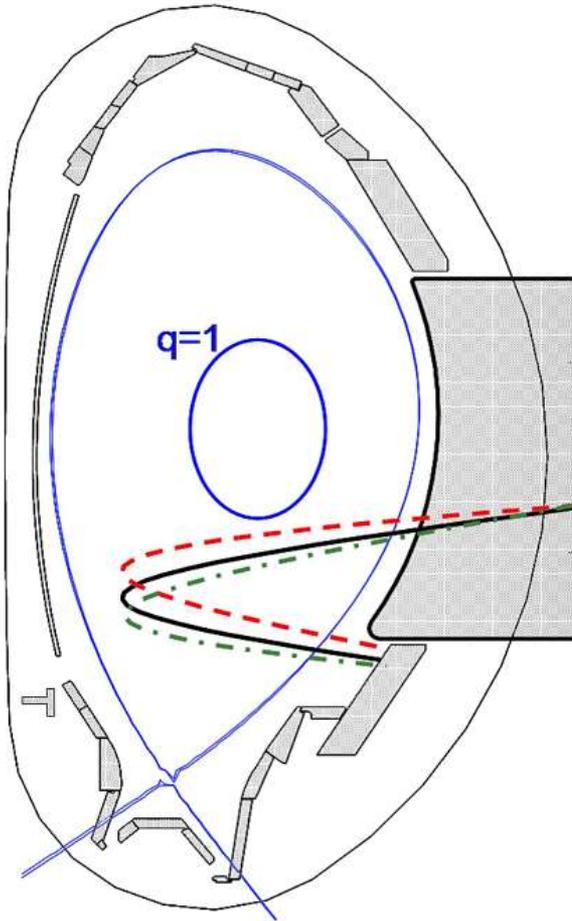
The effect strongly depends from the ^3He concentration.

Thus, the main question is:

How to control the minority concentration in ITER and DEMO?

The strongest effect on mode stability is for a ^3He concentration of only 1%. When there is too much ^3He , the energy of the particles in the tail of the distribution becomes too low to have a strong effect on the kink mode whereas too little ^3He means that the absorbed power is low and the broader distribution function leads to increased fast ion losses.

Changes of the beam direction



[I.T. Chapmen, V.Igochine, et.al.Nucl. Fusion 49 (2009)]
 [I.T. Chapmen et.al. Physics of Plasmas 16 (2009)]

Position of the NBI with respect to $q=1$ surface is important

FIG. 1. (Color online) The beam trajectories of the off-axis PINI in ASDEX Upgrade as the PINI is tilted on its support. Also shown for comparison is the approximate position of the $q=1$ surface.

Changes of the beam direction

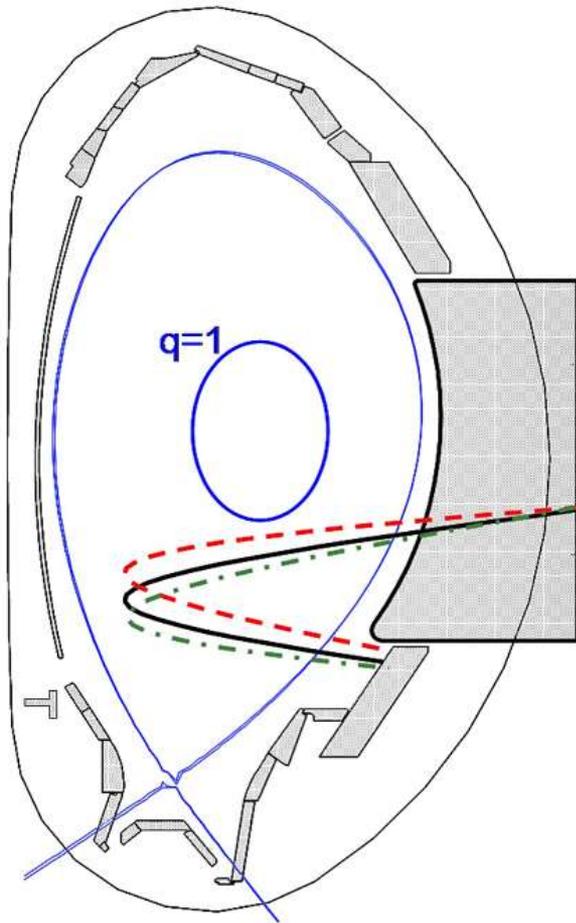
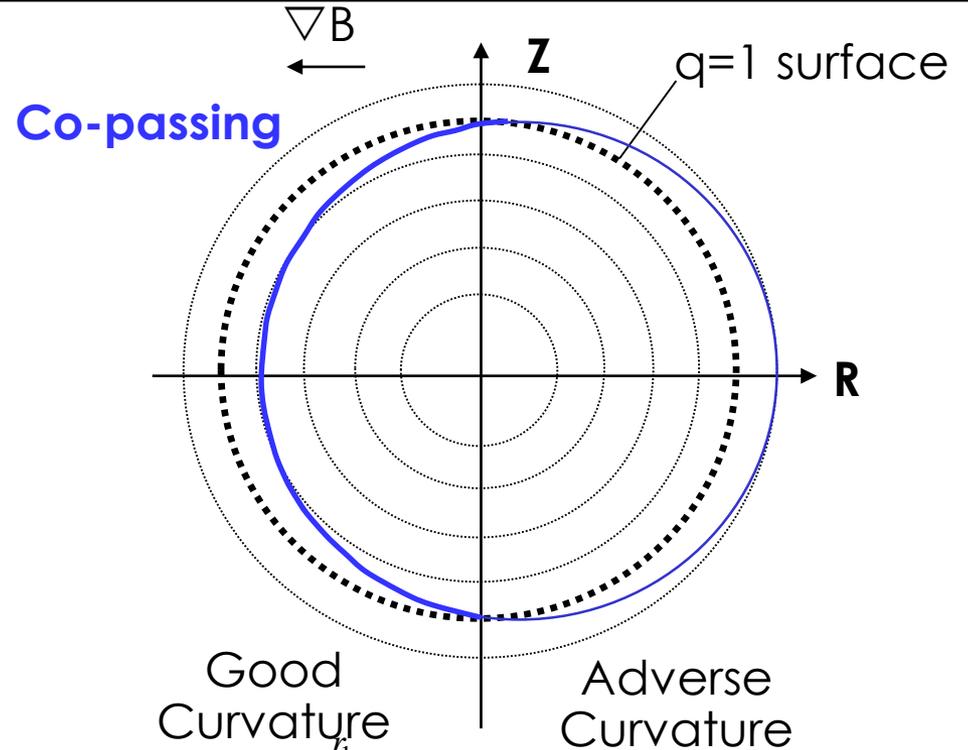


FIG. 1. (Color online) The beam trajectories of the off-axis Pini in ASDEX Upgrade as the Pini is tilted on its support. Also shown for comparison is the approximate position of the $q=1$ surface.



$$\delta W \sim - \int_0^{r_1} (\xi \cdot \nabla \langle P_h \rangle) (\xi \cdot \kappa) dr \neq 0$$

Co-pass, $\langle P_h \rangle' |_{r_1} < 0 \rightarrow$ stabilising

Co-pass, $\langle P_h \rangle' |_{r_1} > 0 \rightarrow$ destabilising

[Graves, PRL, 2004; Chapman et al, PPCF, 2008]

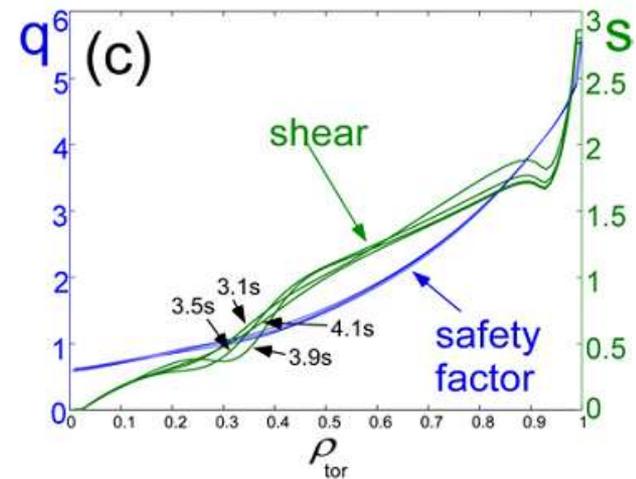
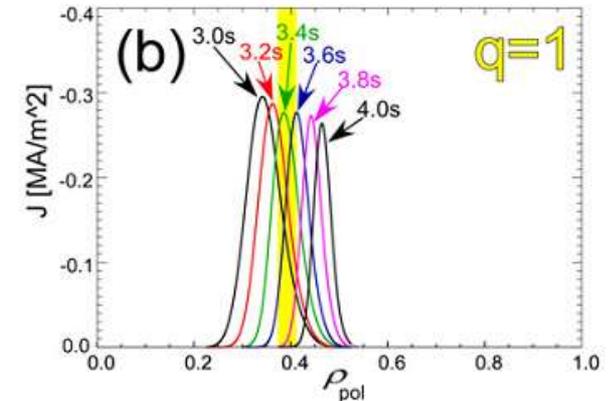
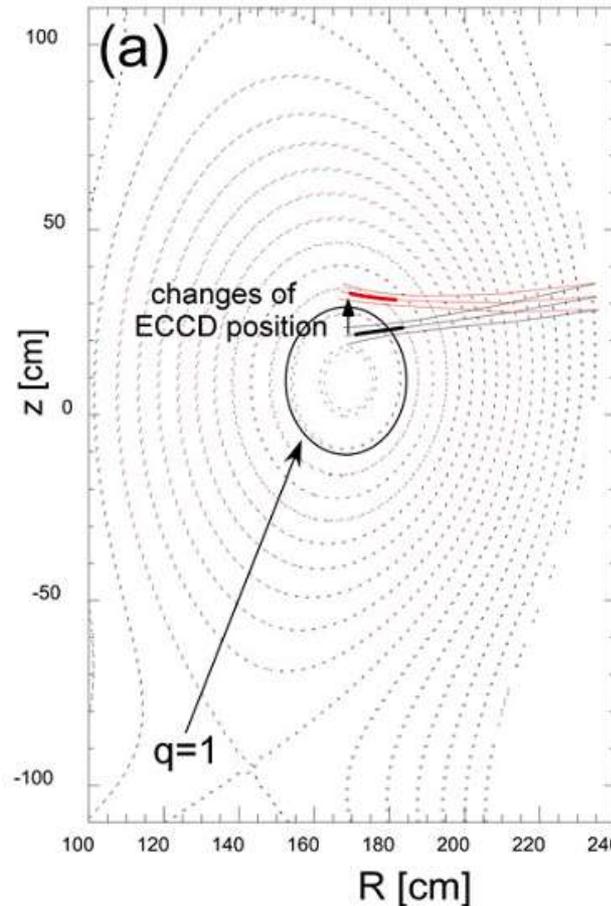
Destabilization of fast particle stabilized sawteeth

ITER relevant situation:

ICRH heating creates fast particles (mock up alpha particles in ITER).

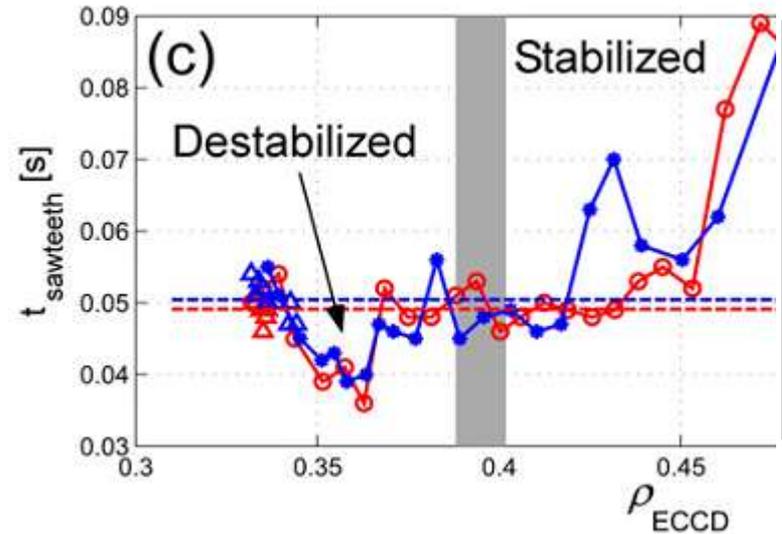
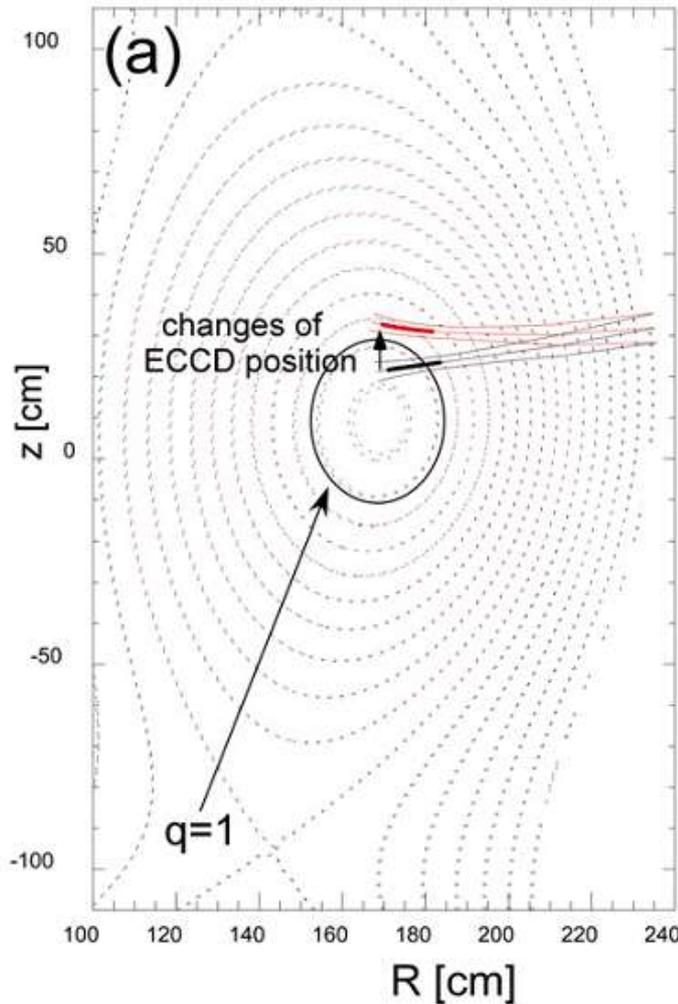
Central ECRH heating keeps the temperature profile constant (this avoids effects T_e profile on the (1,1) mode stability).

ECCD is applied around $q=1$ surface to change shear.



[Igochine et.al. Plasma Phys. Control. Fusion 53 (2011)]

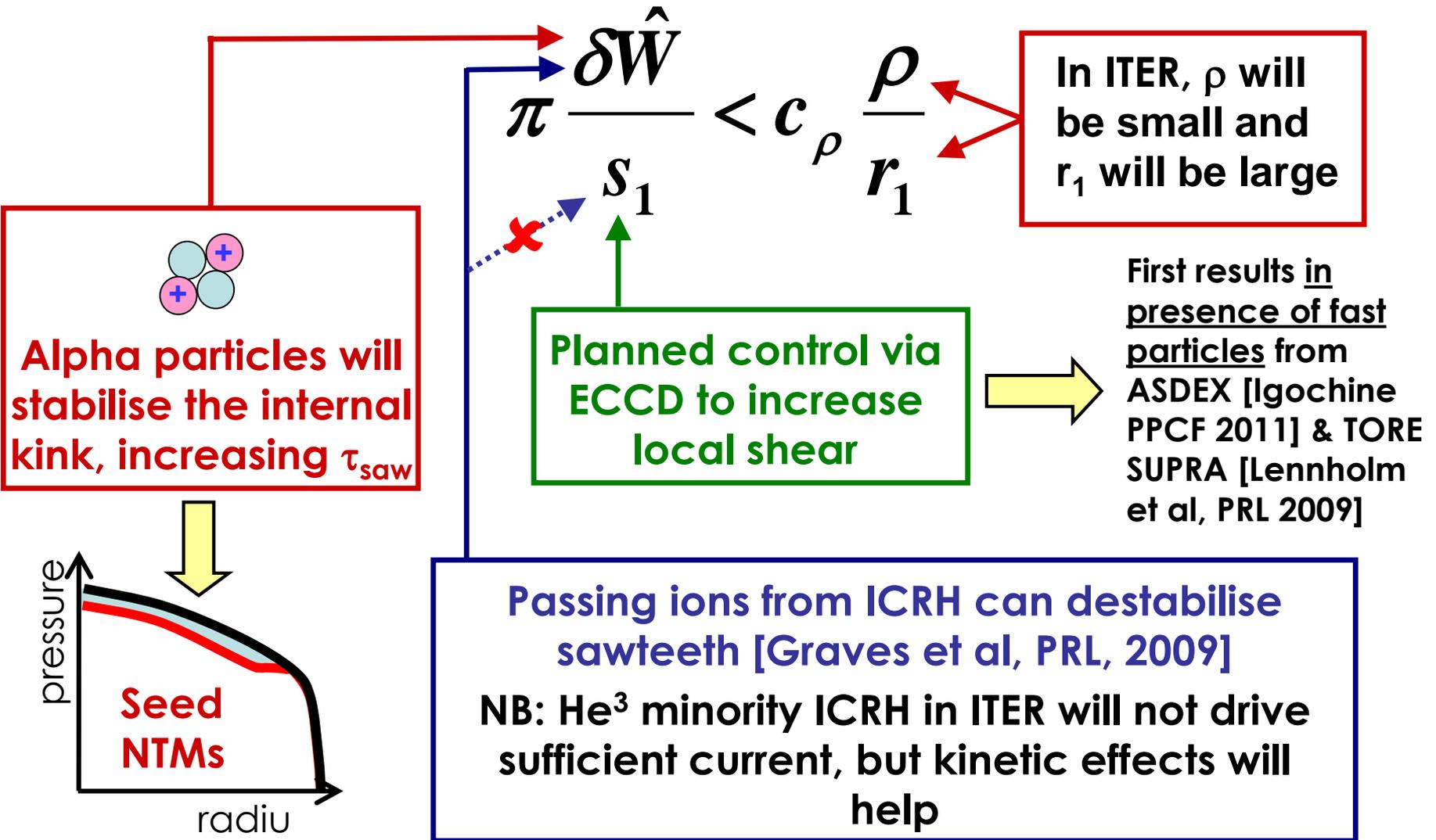
Destabilization of fast particle stabilized sawteeth



About 40% reduction of the sawtooth period is achieved.

Expected sawteeth in ITER is about the critical size to trigger NTMs.

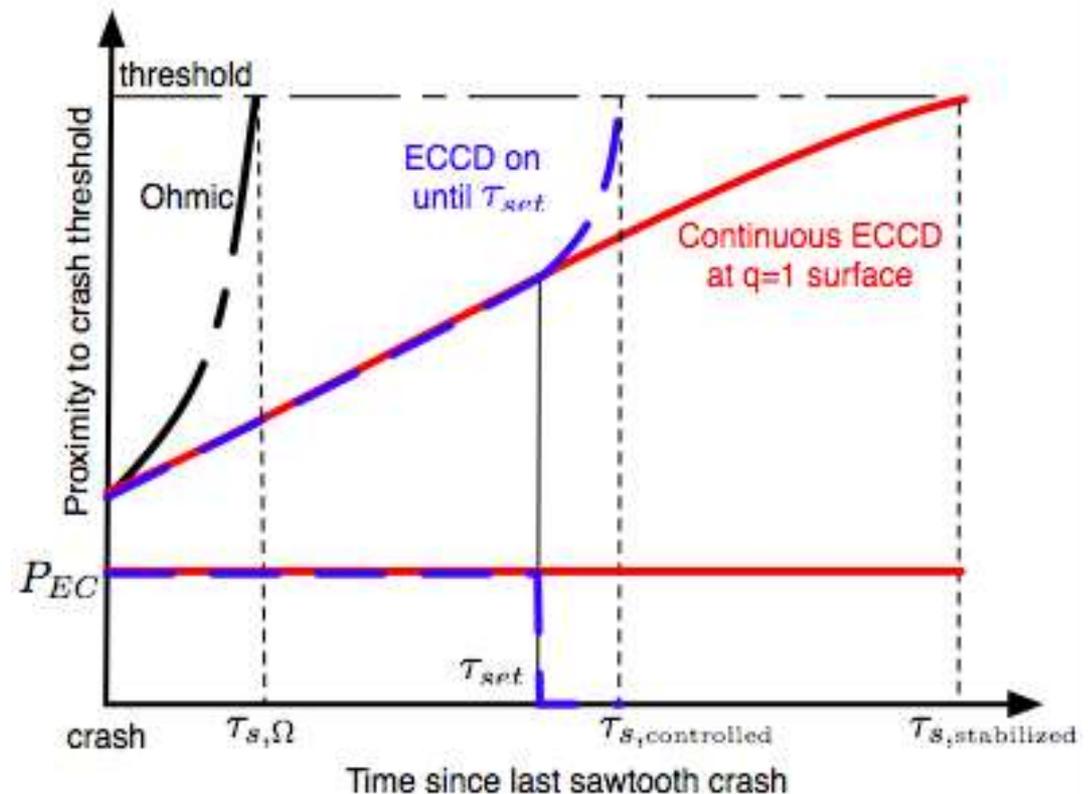
[Igochine et.al. Plasma Phys. Control. Fusion 53 (2011)]



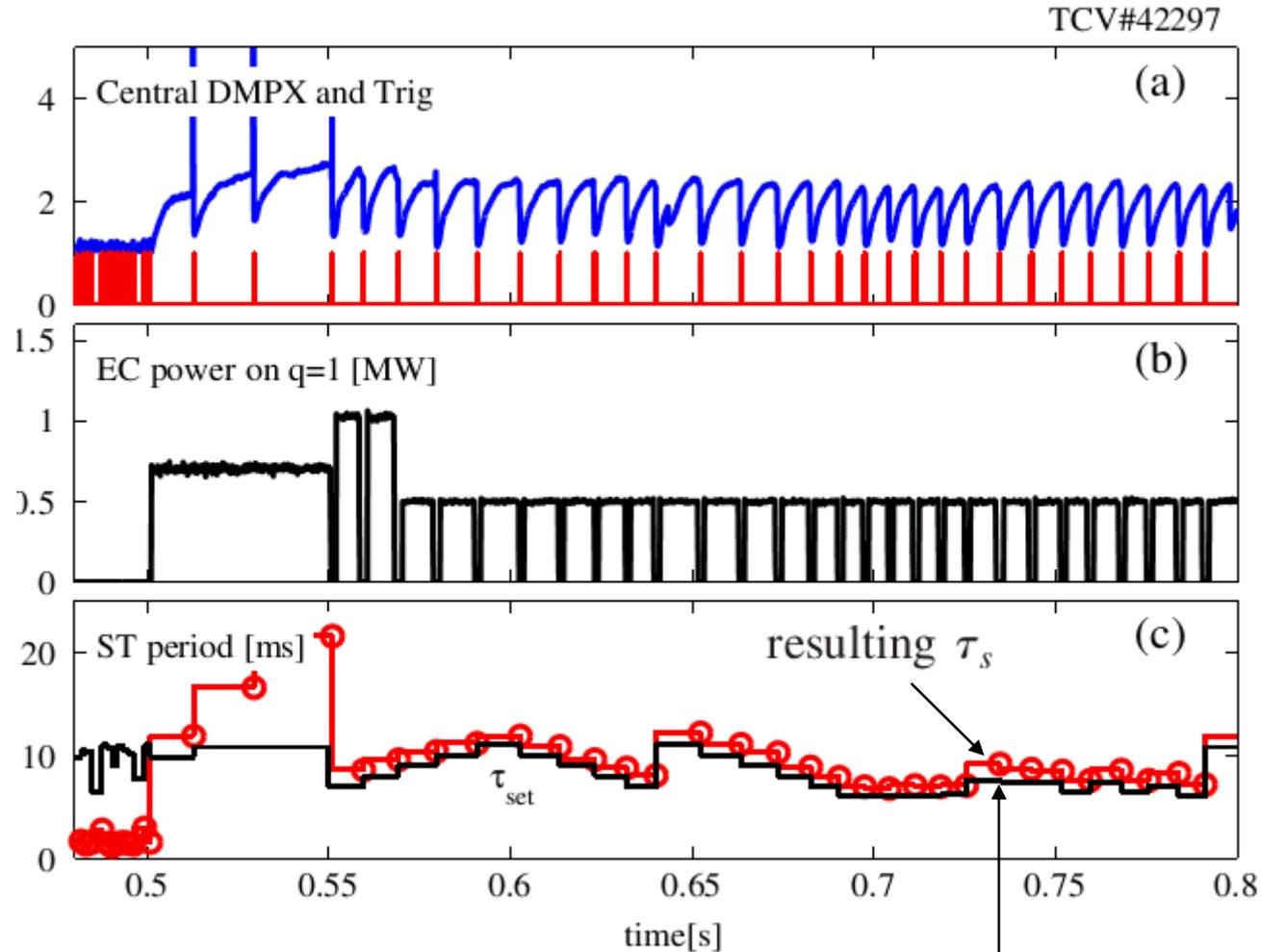
Sawtooth period can be controlled by removing the stabilizing effect during the sawtooth ramp

Switch power of stabilizing actuator OFF after given time

- Crash threshold reached soon after
- Can obtain period in between natural and stabilized period
- TCV case: actuator: ECCD close to $q=1$, threshold: critical shear



Novel methods for sawtooth control, variant 1



T. P. Goodman. PRL **106**, 245002 (2011)

Novel methods for sawtooth control, variant 2

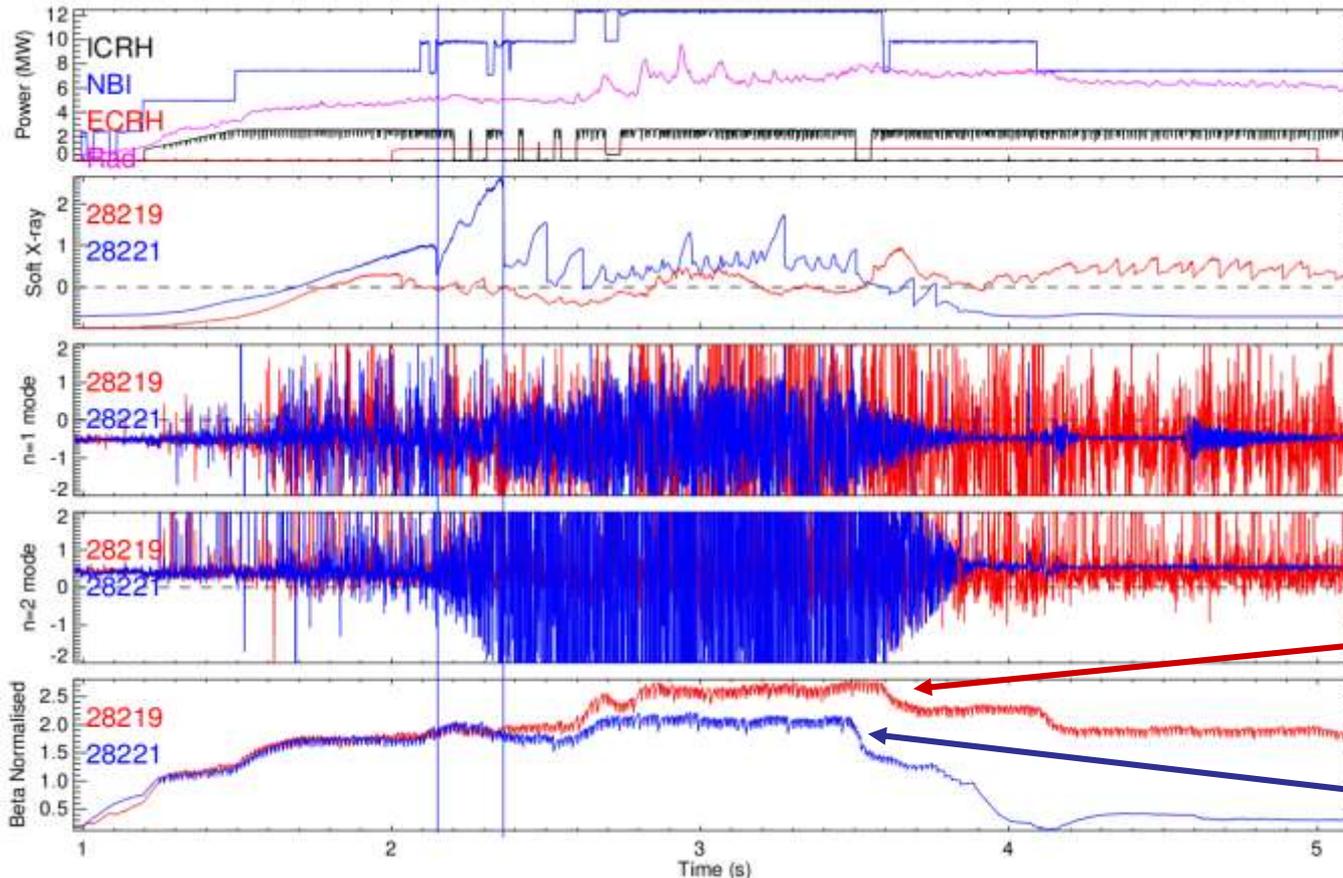


Is a passive action sufficient to avoid NTM?

Yes, if we drive small current inside $q=1$.

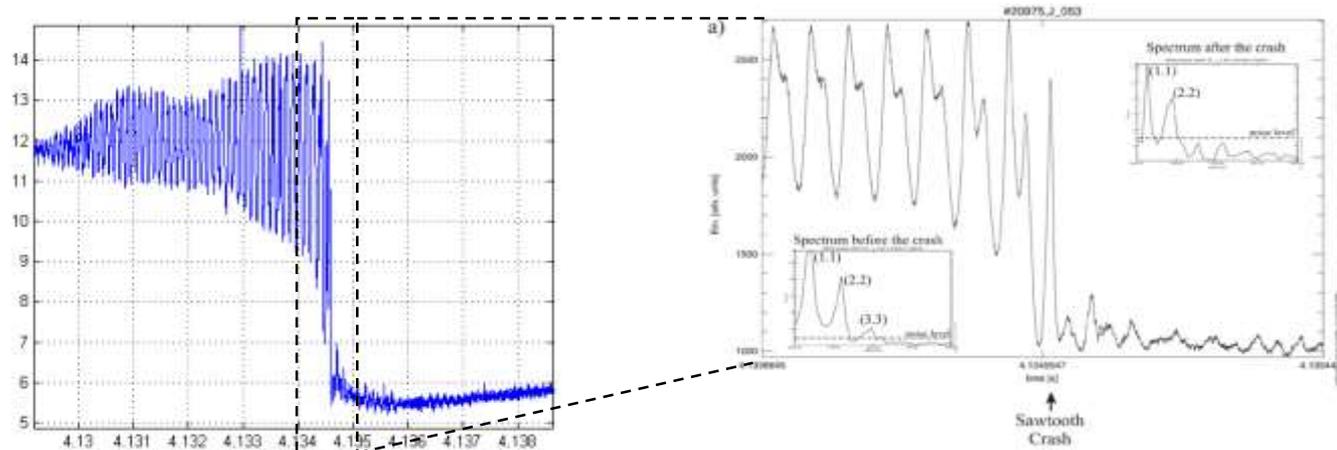
(No precise position is required in contrast to NTM case!)

ASDEX Upgrade
Recent results (2012)



No NTM with
1MW gyrotron

NTM drops β_N
by 25%!



Understanding of the sawteeth

Slow time scale

- sawtooth period
- mixing radius
- amplitude

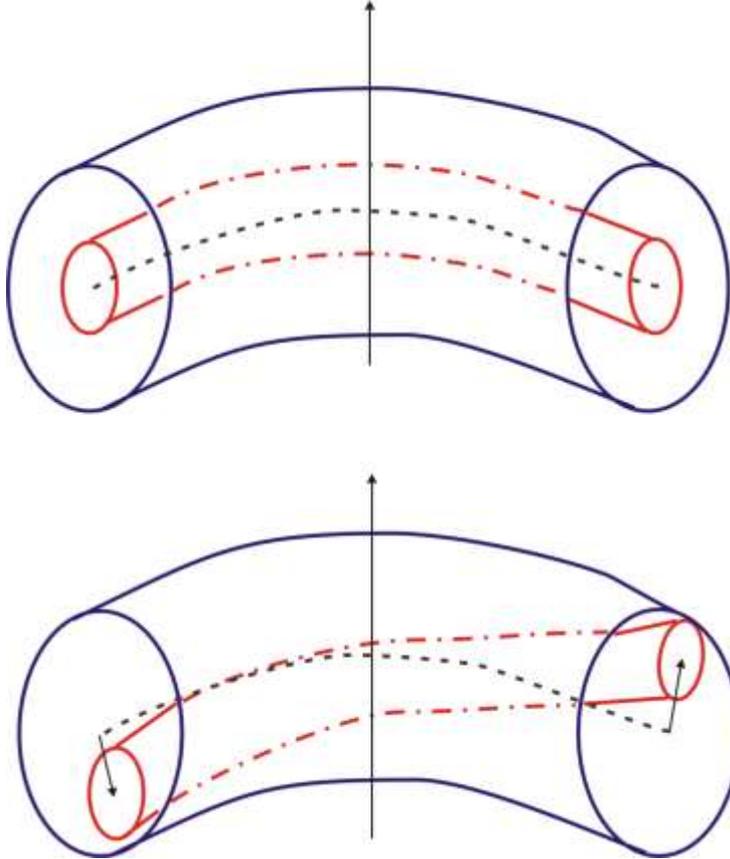
Especially important for ITER

Fast time scale

- mechanism of the crash

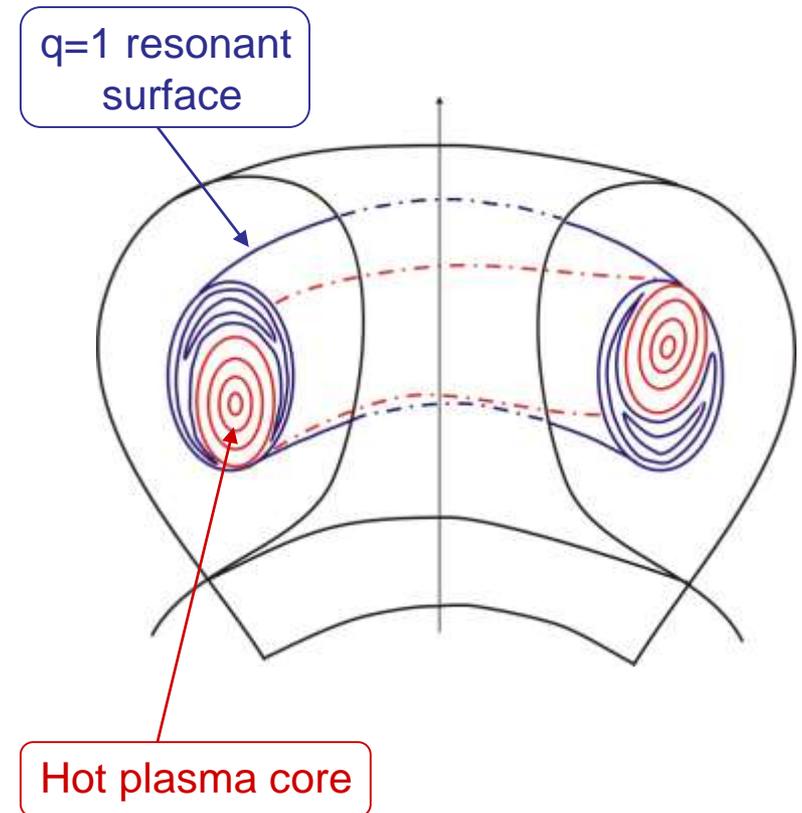
Understanding of the basic physics

Internal Kink instability in a Tokamak.



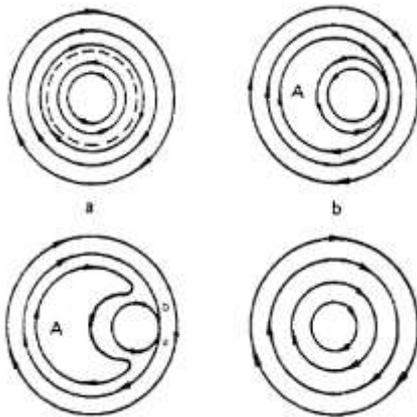
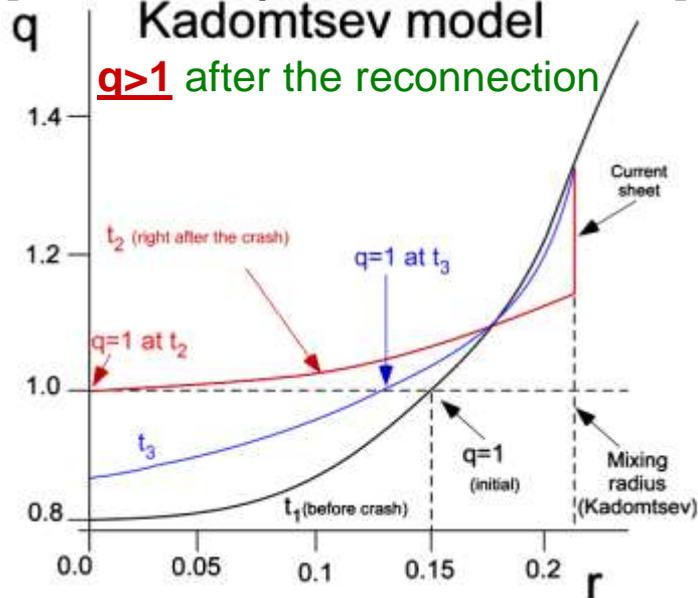
A Tilt and Shift of the Core Plasma.

Sawteeth: internal (1,1) kink mode.



Sawtooth crash (fast time scale)

[Кадомцев физика плазмы 1975]

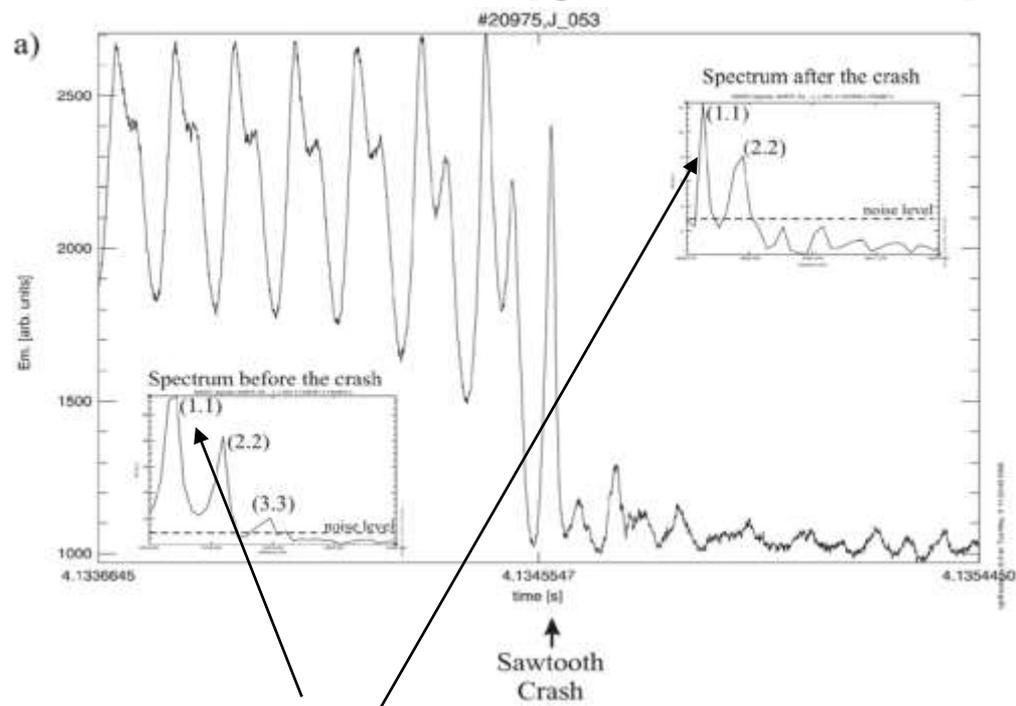


O-point becomes the new plasma center

ASDEX Upgrade discharge

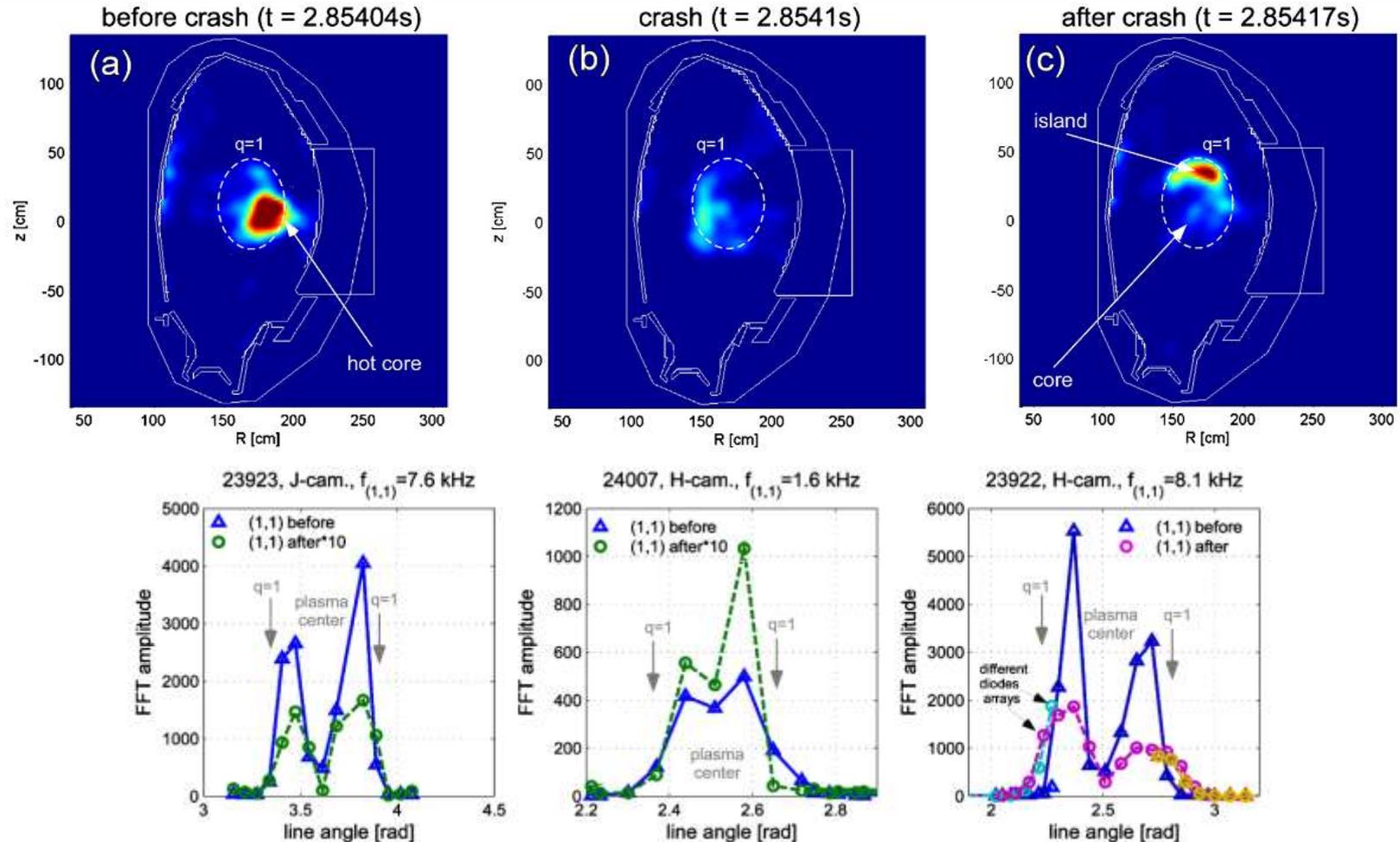
$q < 1$ after the reconnection

[Igochine et. al., NF 2007]



(1,1) mode before and after the crash displays existence of $q < 1$ in the plasma after the crash

Sawtooth crash (fast time scale)

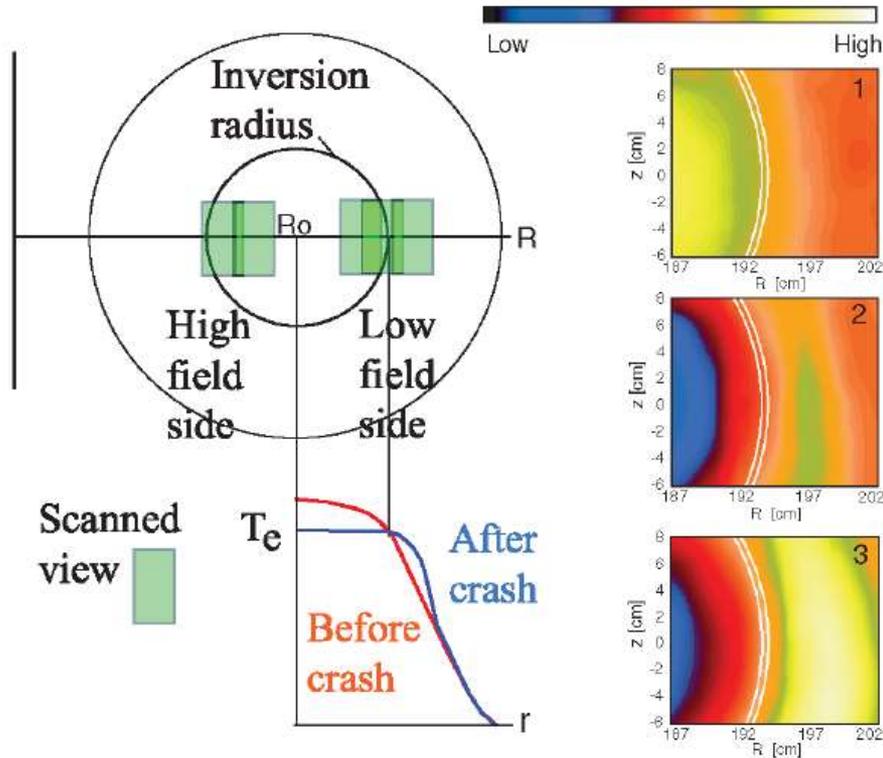


Position of (1,1) mode is the same before and after the crash!

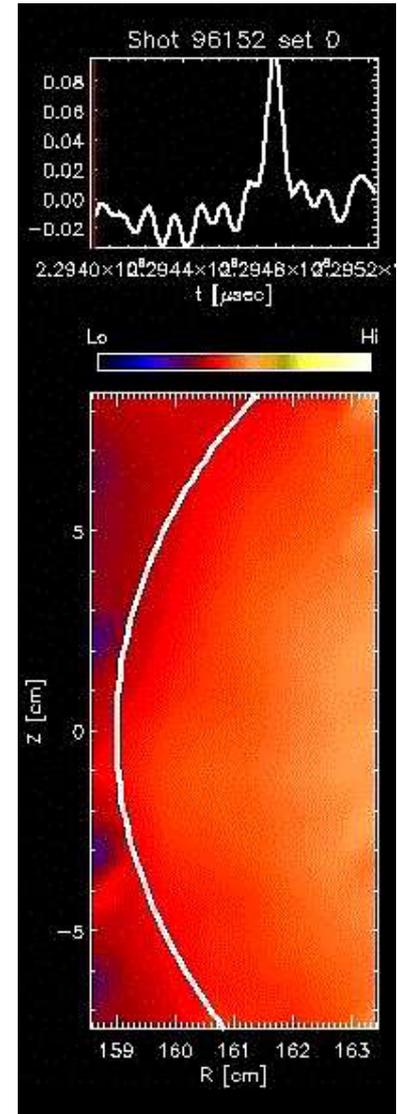
[Igochine et.al. Phys. Plasmas 17 (2010)]

Sawtooth crash in TEXTOR

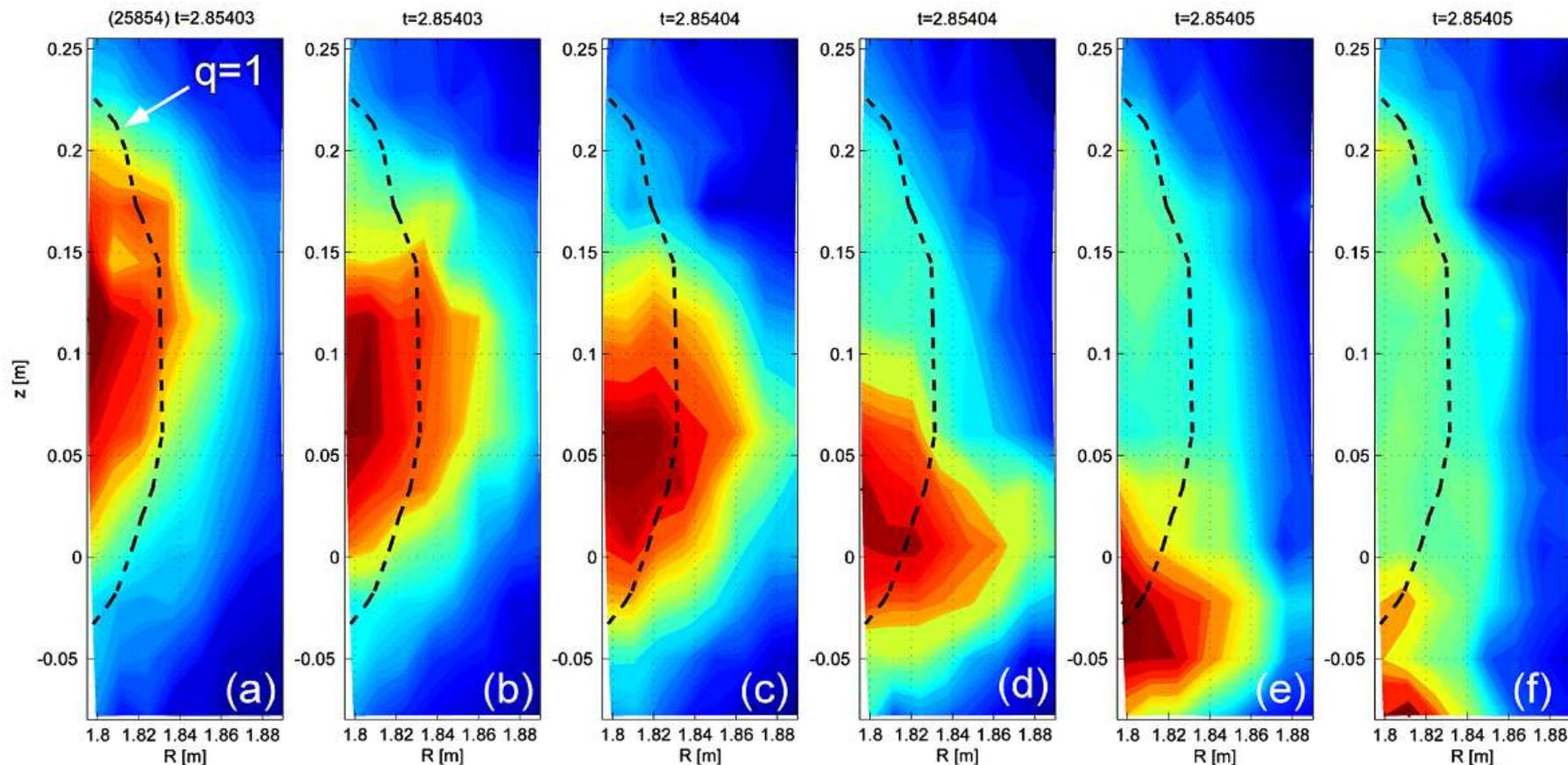
High field side reconnection in TEXTOR [Park, PRL and PoP 2006]



1. High field side reconnection
2. Poloidally localized reconnection



Sawtooth crash in ASDEX Upgrade



Local outflow of the hot core through the X-point. The same as in TEXTOR.

[Igochine et.al. Phys. Plasmas 17, (2010)]

The sawtooth crash model has to fulfilled the following conditions:

1. (1,1) mode remains after the crash
2. Position of the mode is not affected by the crash
3. Temperature from the core is removed
4. Heat outflow is local
5. Heat outflow is fast
6. Reconnection could happened in any poloidal locations, also on the high field side

Kadomtsev model [Кадо́мцев физика плазмы1975] contradicts to 1,2,5

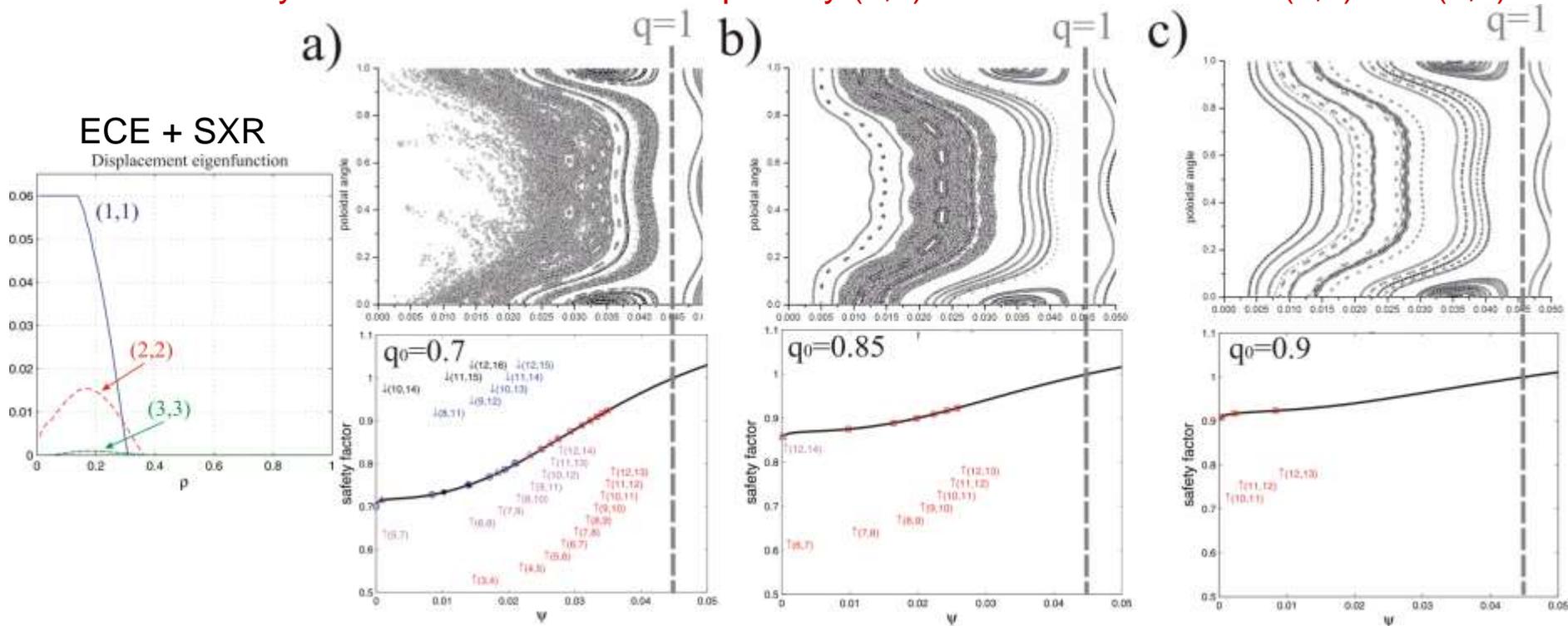
Balooning model [Nishimura PoP 1999] contradicts to 6

....

Stochastic model could be a possible explanation

Main idea: interaction of the modes leads to stochastisity which removes heat from the core and keep the mode at its original position.

Stochasticity is a result of interaction of primary (1,1) mode with harmonics (2,2) and (3,3)

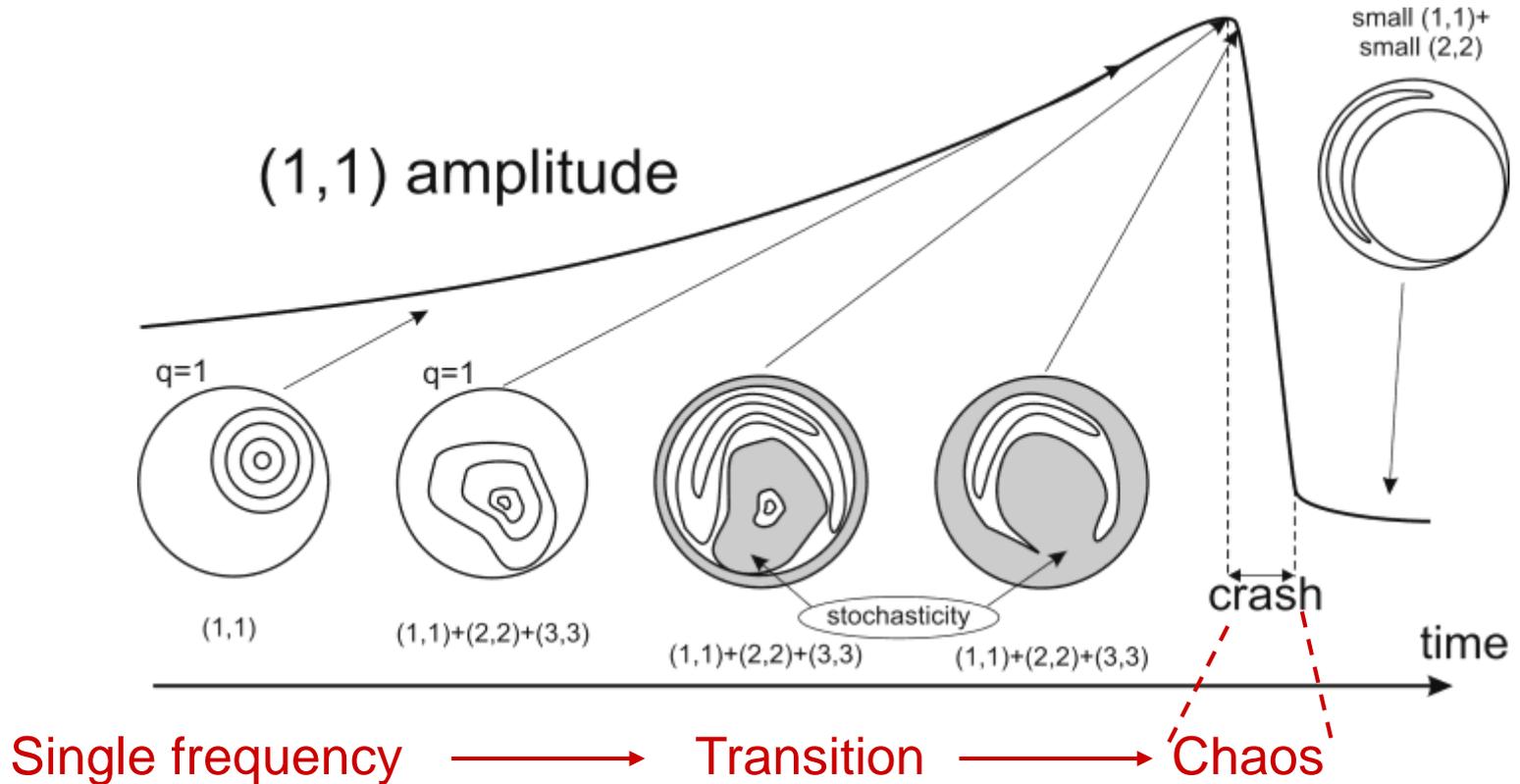


Conclusions:

- Experimental perturbations are sufficient to stochastize the region
- Stochastization requires the existence of several low-order rational surfaces
- q_0 determines the number of such surfaces → **the main factor is q_0 !**

[Igochine et. al., NF 2007]

Field line tracing assumes static problem. All perturbations are coupled by definition. What could we say about dynamics of the process?



Hilborn "Chaos and Nonlinear Dynamics"

Intermittency

Quasi-periodicity

Roads to chaos

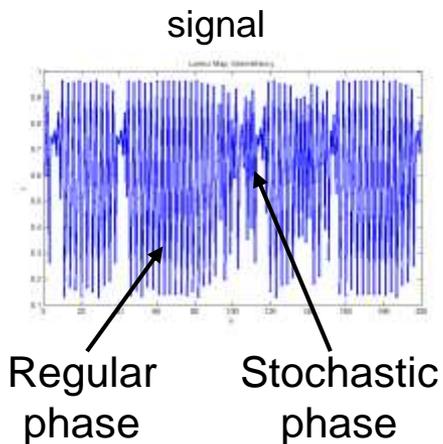
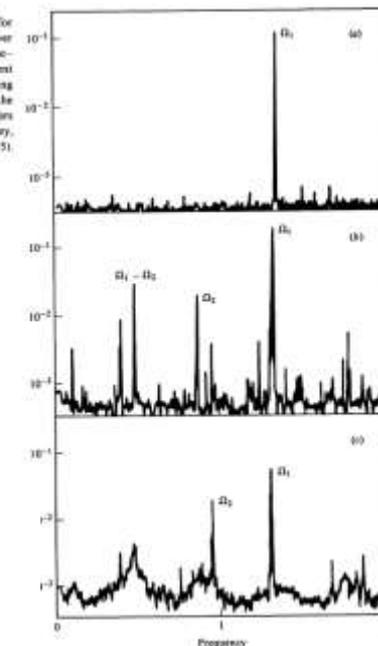
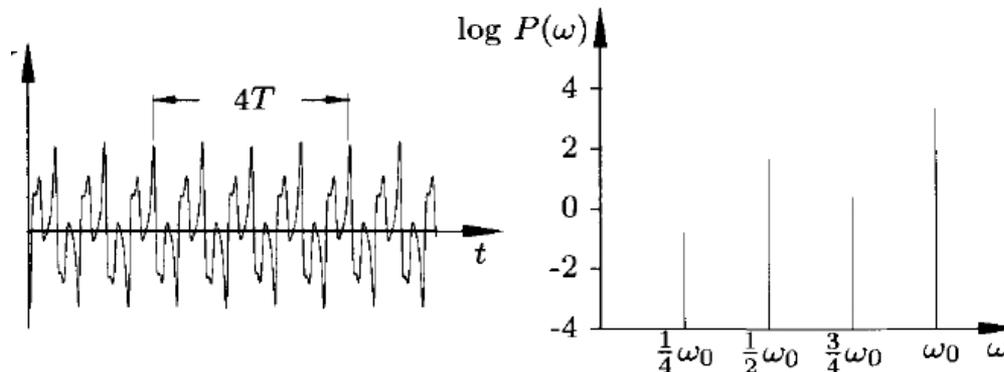


Figure 6.1 Results for frequency power spectra for a Couette-Taylor experiment with increasing rotation rate of the inner cylinder (Gollub and Swinney, 1975)



Period doubling



Physical, chemical, biological systems use one of these variants.
Roads to chaos are universal!

Hilborn "Chaos and Nonlinear Dynamics"

Intermittency

Quasi-periodicity

Roads to chaos

Period doubling

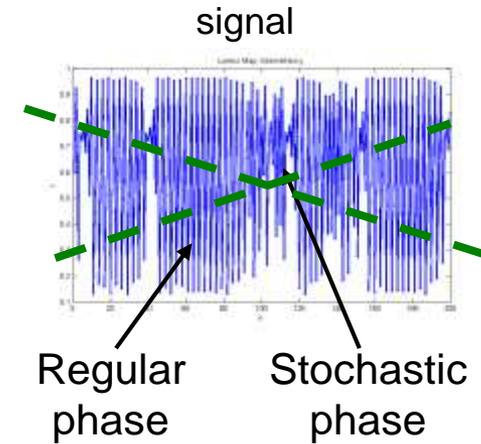
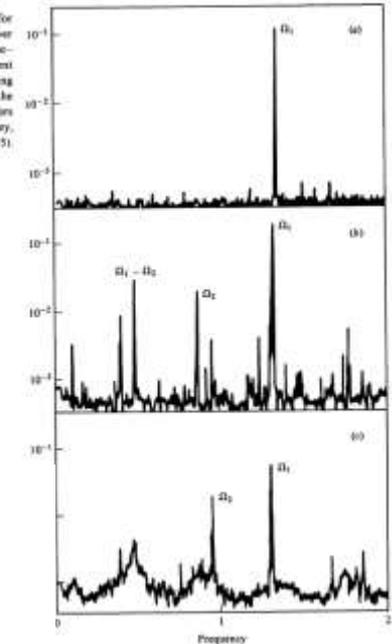
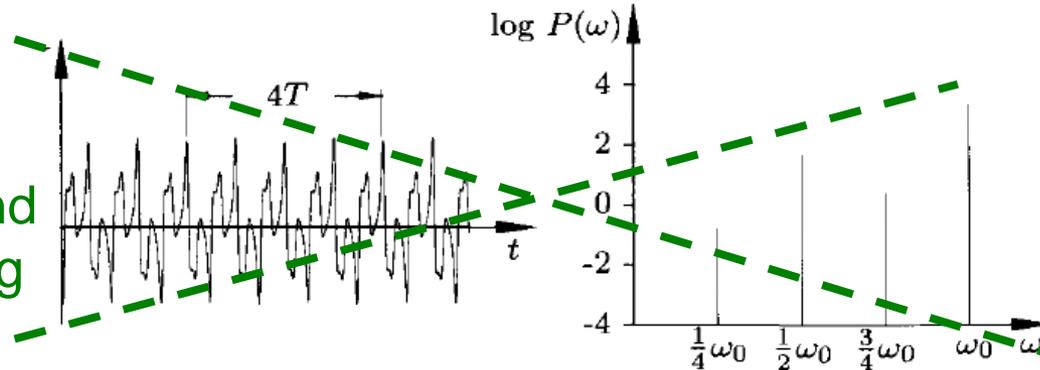


Figure 6.1 Results for frequency power spectra for a Couette-Taylor experiment with increasing rotation rate of the inner cylinder (Gollub and Swinney, 1975)



Sawtooth show no intermittency and no period doubling



Possible candidate is only quasiperiodicity

Observation of quasiperiodic transition in a crystal

Single freq. with harmonics

Quasiperiodic phase

Chaos

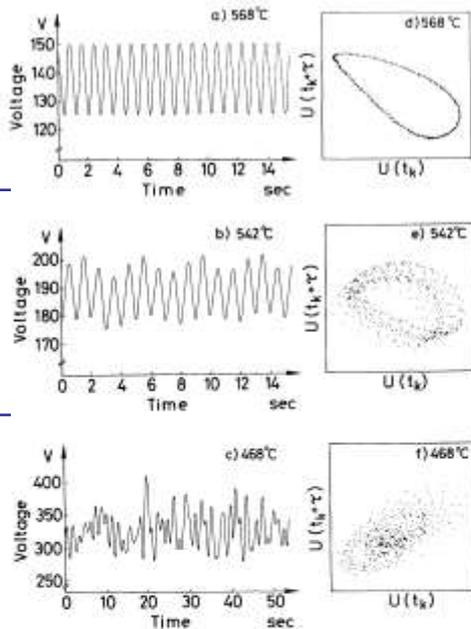


FIG. 1. Self-generated oscillations of the voltage $U(t)$ and corresponding phase portraits $U(t_k + \tau)$ vs $U(t_k)$ ($k=1-500$ is the index for the sampling points; $\tau=0.3$ sec is the time delay). The temperature is varied; the current density (1.8 mA/cm^2) and oxygen flow rate (1 L/h) are held constant. Note the change of scale in (c).

irregular pattern [Fig. 1(f)].

Simultaneous with the oscillations, the birefringence pattern of the crystal is seen to be locally disturbed: A "domain" emerges from the cathode as the voltage increases to the maximum, and disperses gradually during its movement through the crystal as the voltage relaxes to the minimum. Measurements of the voltage across each of three successive sections of the crystal, made with use of intermediate platinum wire loop electrodes, show

[S.Martin et al., PRL, 1984]

Single freq. with harmonics

Quasiperiodic phase.

Two incommensurable frequencies and their linear combinations

Chaos

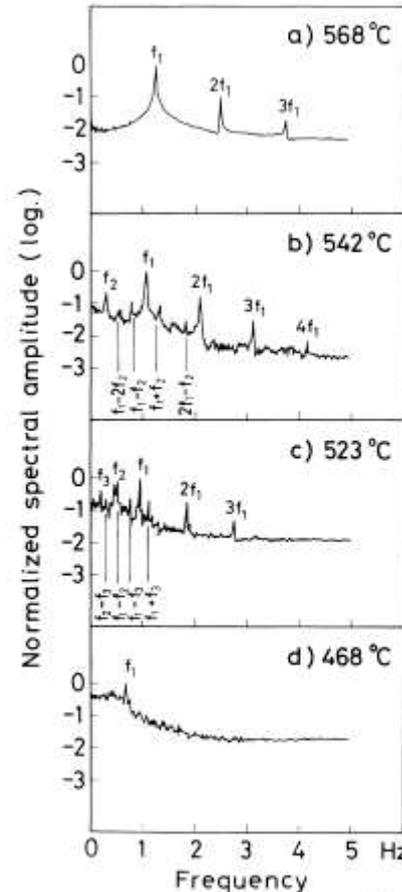
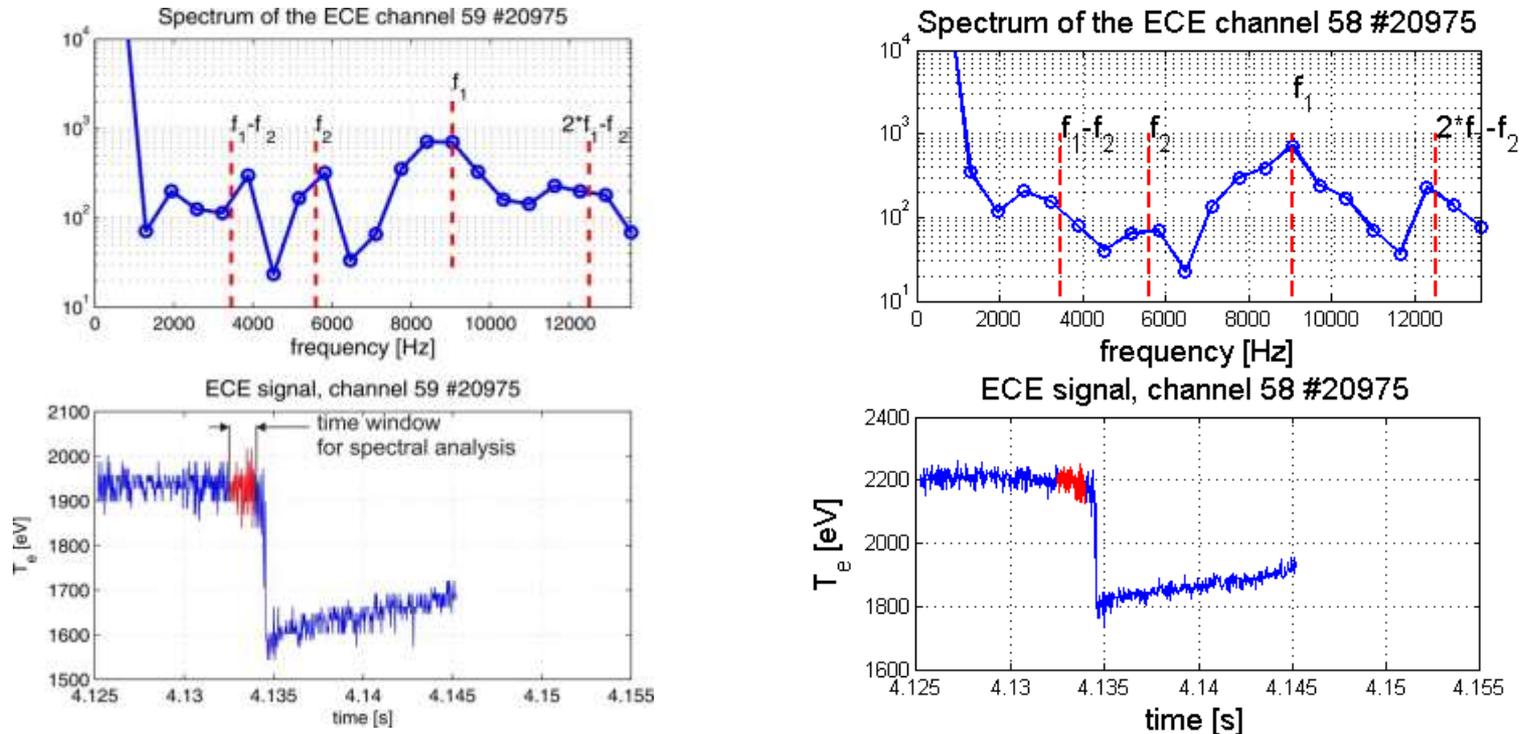


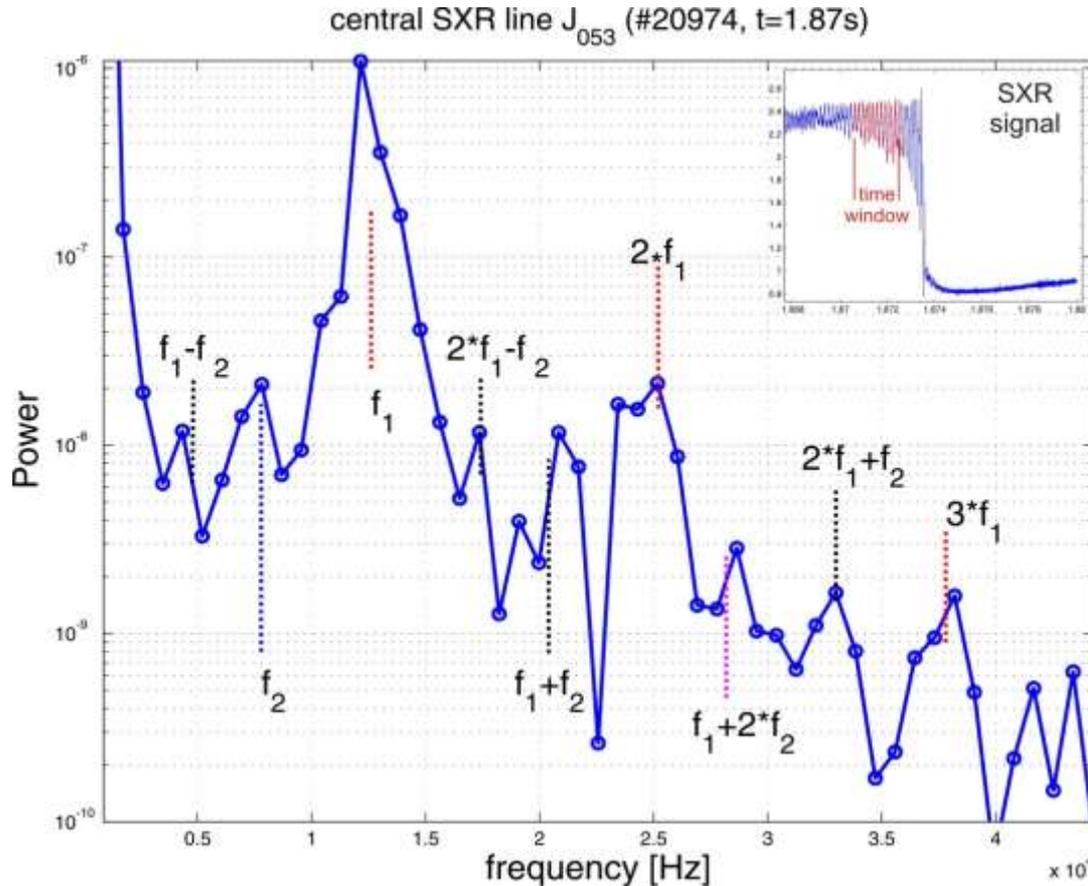
FIG. 2. Normalized spectra of $U(t)$ in the (a) oscillatory, (b),(c) intermediate, and (d) chaotic states. The spectral amplitude is normalized in relation to the maximum of the highest peak. In (b) and (c) two and three fundamental frequencies, respectively, are seen.

ECE power spectrum and the signal for the same sawtooth crash from two central ECE channels



These measurements resolve also other combinations of the resonances which are not seen in SXR signals

Transition to chaos from central SXR signal



Other discharge show the same spectrum in slightly quasiperiodic stage.

The low frequency spectrum is completely described by two primary frequencies and their linear combinations.

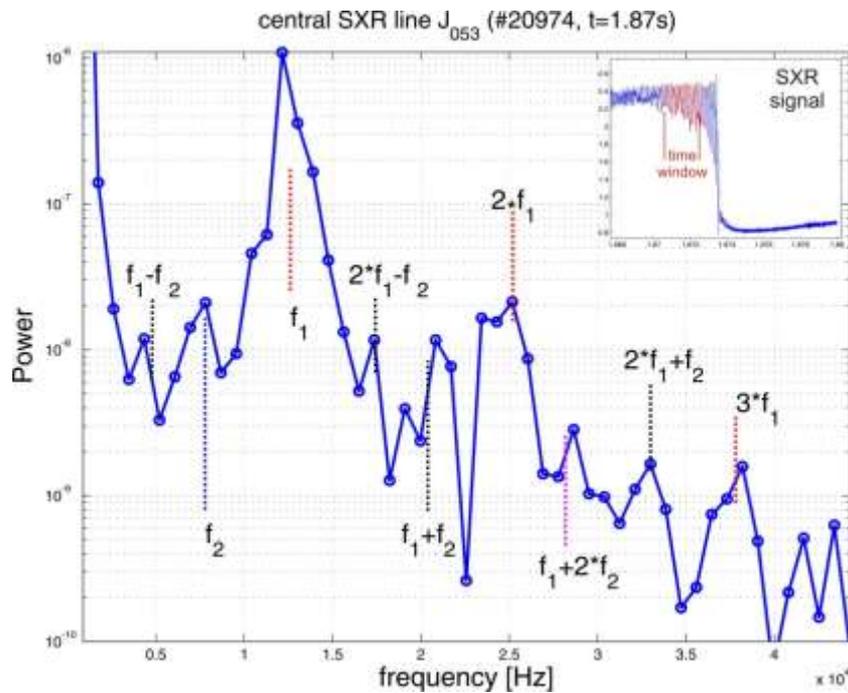
Relation between two frequencies is golden mean ("the most irrational number")

$$\frac{f_2}{f_1} = \frac{\sqrt{5}-1}{2} \approx 0.618$$

This fact indicates that the chaos in the system is approached in the most "intense" way

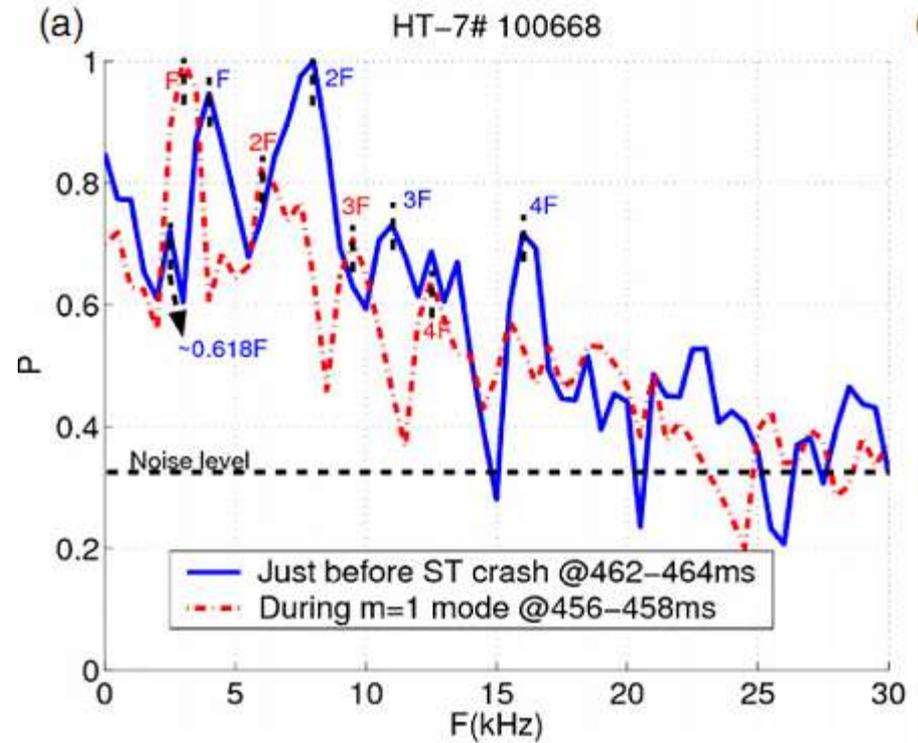
[Igochine et.al. Nucl. Fusion 48 (2008)]

ASDEX Upgrade

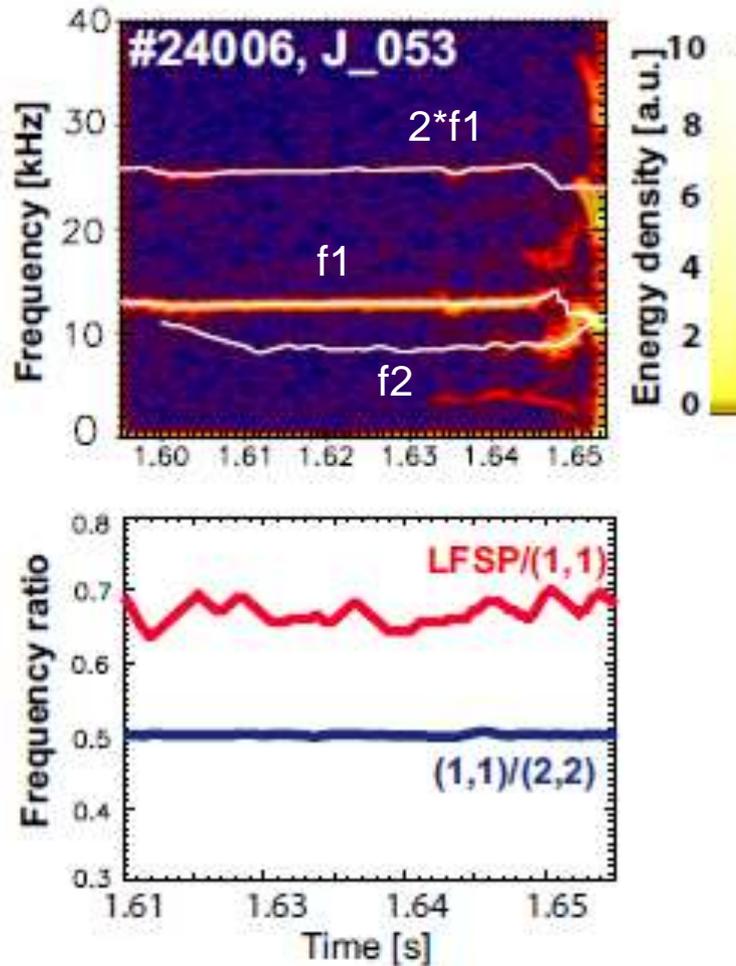


[Igochine et.al. Nucl. Fusion 48 (2008)]

HT-7



[Sun et.al. Plasma Phys. Control. Fusion 51 (2009)]



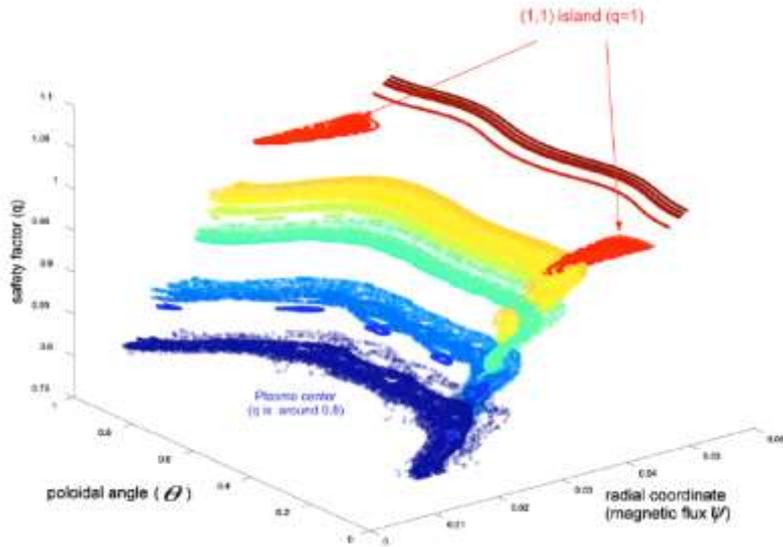
More detail investigation of a large set of sawtooth crashes (wavelet technique, coherence technique and bi-coherence technique) :

- Confirm existence of the mode with golden mean relation to the primary (1,1)
- Give out typically (1,1) mode structure for this mode as well
- Long precursor phase is required to identify the second mode!

[Papp et.al. Plasma Phys. Control. Fusion 53 (2011)]

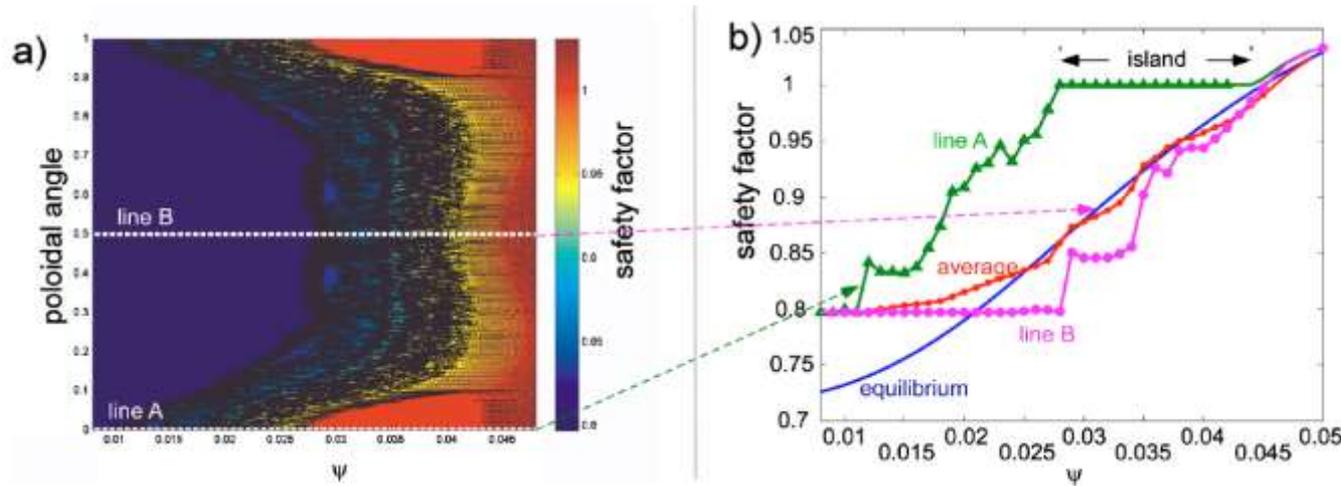
Changes of the safety factor during the crash in stochastic model

[Igochine et.al. Phys. Plasmas 17 (2010)]



$$q = \lim_{\Delta\phi \rightarrow \infty} \frac{\Sigma \Delta\phi}{\Sigma \Delta\theta} \approx \frac{2\pi \cdot (N-1)}{\Sigma_{i=1}^{N-1} (\theta_{i+1} - \theta_i)}$$

Safety factor remains almost the same in stochastic model (contrary to Kadomtsev model!)

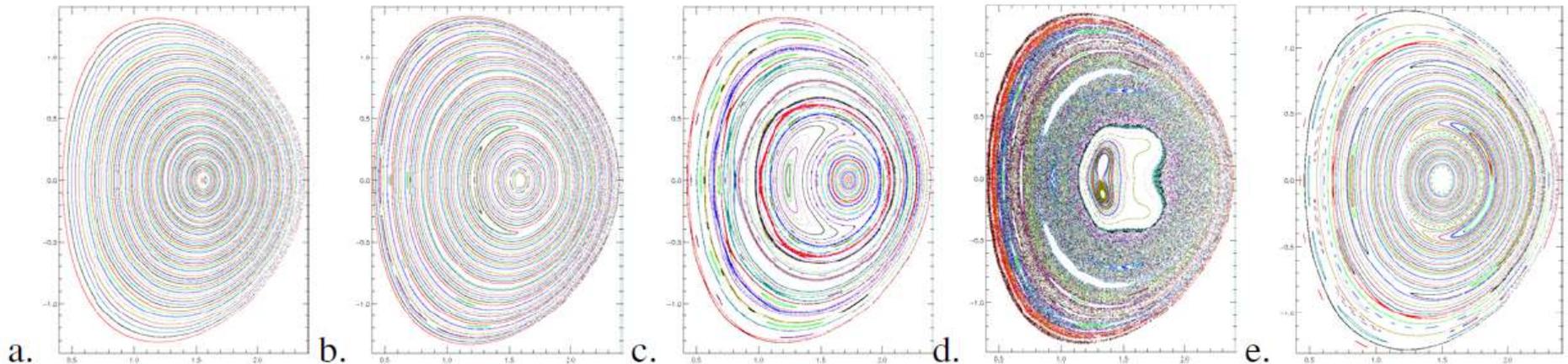


Sawtooth modelling

- Nonlinear MHD simulations (M3D code) show stochasticity.
- but .. „multiple time and space scales associated with the reconnection layer and growth time make this an extremely challenging computational problem. ... and there still remain some resolution issues.”

Small tokamak → small Lundquist number: $S = 10^4$ (big tokamaks 10^8)

Lundquist number = (resistive diffusion time)/(Alfven transit time)



[Breslau et.al. Phys. Plasmas 14, 056105, 2007]

**Non-linear simulations of the sawtooth is very challenging task
(even in a small tokamak).**

Sawtooth modelling

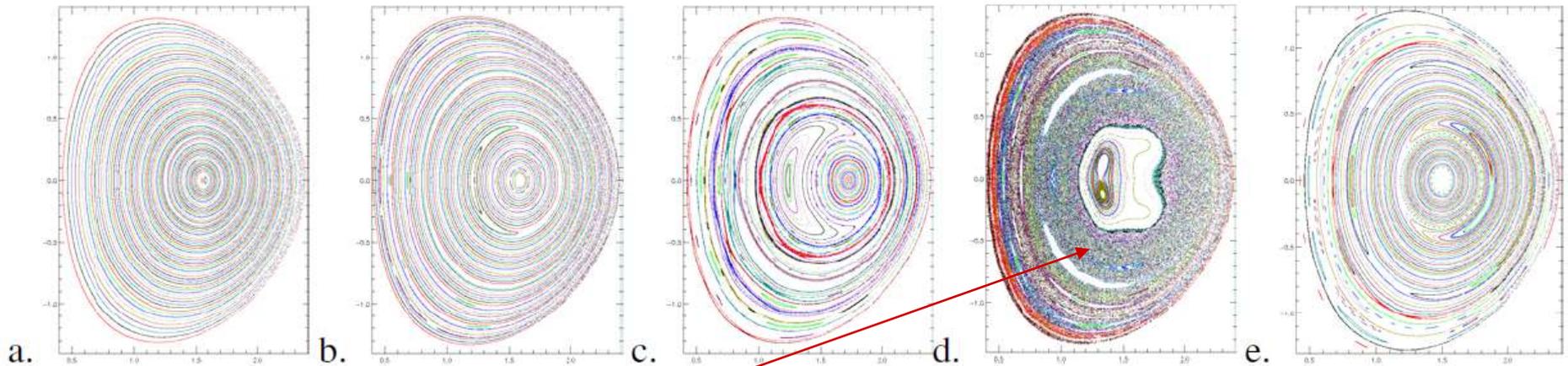
Ohm's law, 2 fluid MHD

$$\eta \vec{j} = \underbrace{\vec{E} + \vec{v} \times \vec{B}}_{\text{ideal MHD}} - \underbrace{\frac{1}{en_e} \vec{j} \times \vec{B}}_{\text{Hall term}} + \underbrace{\frac{1}{en_e} \nabla p_e}_{\text{electron pressure term}}$$

resistive MHD

...at least two fluid MHD with correct electron pressure description are necessary for reconnection region (fast crash time, smaller stochastic region)!

We will discuss this point in details in the last lecture



[Breslau et.al. Phys. Plasmas 14, 056105, 2007]

Stochastic region is too large, ... much more than visible in the experiments (heat outflow is rather global instead of local as in the experiments)

Stochastic model explains:

- Constant position of the $q=1$ surface during the crash (1,1)
- Fast remove of the heat from the core
- It has no contradiction with observations
- Indications of the transition to stochastic phase are found

MHD calculations show stochasticity during the crash... but the degree of stochastization is too big (not all important physics is inside)

Degree of stochastization is not clear. Two situations are possible:

- Fully stochastic core
- Partially stochastic core (only along the separatrix)

In both cases, (1,1) island is not destroyed in stochastic model!

We are quite successful in controlling of sawteeth even with fast particles (ITER relevant situation). Many control techniques are developed.

During the last years a new important information about the crash phase were obtained

Stochastic model is a possible candidate for the explanation of the sawtooth crash.

There are still lack of sawtooth modeling with full nonlinear MHD codes with 2 fluid effects.

Full understanding of the crash phase is necessary for exact predictions of sawtooth amplitude and period for ITER and DEMO.

The physics of sawtooth stabilization

I T Chapman¹, S D Pinches¹, J P Graves², R J Akers¹, L C Appel¹,
R V Budny³, S Coda², N J Conway¹, M de Bock⁴, L-G Eriksson⁵,
R J Hastie¹, T C Hender¹, G T A Huysmans⁵, T Johnson⁶,
H R Koslowski⁷, A Krämer-Flecken⁷, M Lennholm⁵, Y Liang⁷,
S Saarelma¹, S E Sharapov¹, I Voitsekhovitch¹, the MAST and TEXTOR
Teams and JET EFDA Contributors⁸

Plasma Phys. Control. Fusion **49** (2007) B385–B394

SAWTOOTH INSTABILITY IN TOKAMAK PLASMAS

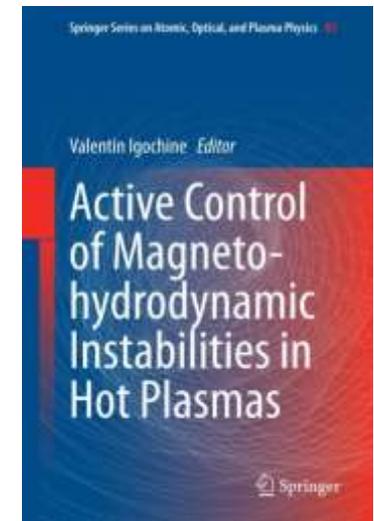
R. J. HASTIE

Imperial College, London, UK

Astrophysics and Space Science **256**: 177–204, 1998.

© 1998 Kluwer Academic Publishers. Printed in the Netherlands.

Chapter 4 of this
book, written by
I.Chapman



V. Igochine, “Active Control of Magneto-hydrodynamic Instabilities in Hot Plasmas”, Springer Series on Atomic, Optical, and Plasma Physics, Vol. **83**, 2015